

# Experimental techniques to investigate the $p$ process of nucleosynthesis

- The „exotic“  $p$  process
- „Direct“ measurements with real and virtual photons
- „Indirect“ measurements:  
Radiative capture



*Andreas Zilges*  
Institut für Kernphysik  
Universität zu Köln

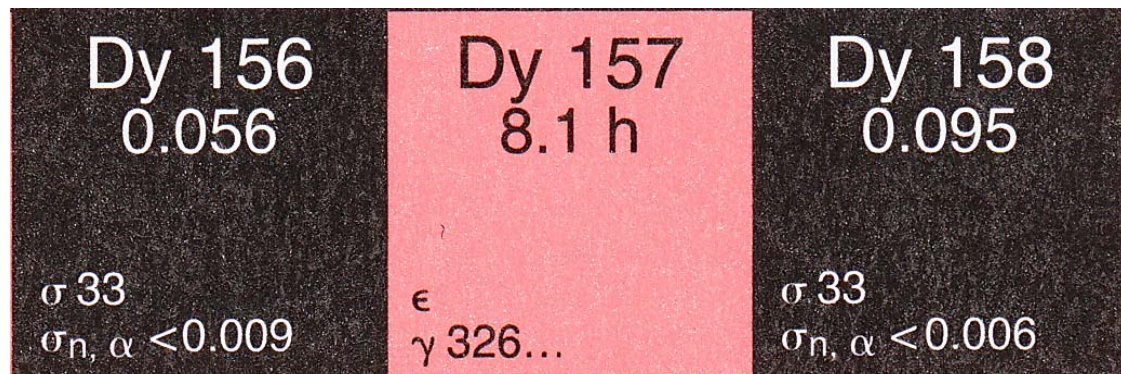
# p nuclei

Er 156 18.6 m $\epsilon; \beta^+ \dots$ $\gamma 35; 30 \dots$ $e^-$	Er 157 18.65 m $\epsilon; \beta^+ 2.5$ $\gamma 53; 391;$ 121...	Er 158 2.25 h $\epsilon$ $\beta^+ 0.8 \dots$ $\gamma 72; 387 \dots$ $m_1$	Er 159 36 m $\epsilon$ $\beta^+ 1.1 \dots$ $\gamma 624; 649 \dots$ $g; m$	Er 160 28.6 h $\epsilon$ $\gamma 7; e^-$ $m_1$	Er 161 3.24 h $\epsilon$ $\beta^+ \dots$ $\gamma 827 \dots$ $g; m$	Er 162 0.139 $\sigma 19$ $\sigma_n, \alpha < 0.011$	Er 163 175 m $\epsilon$ $\beta^+ \dots$ $\gamma (114 \dots)$	Er 164 1.601 $\sigma 13$ $\sigma_n, \alpha < 0.0012$	Er 165 10.3 h $\epsilon$ $\beta^+ \dots$ $\gamma$	Er 166 33.503 $\sigma 3 + 14$ $\sigma_n, \alpha < 7E-5$	Er 167 2.3 s 22.869 $\gamma 208$ $e^-$ $\sigma 650$ $\sigma_n, \alpha 3E-6$
Ho 155 48 m $\epsilon$ $\beta^+ 1.8$ $\gamma 240; 136 \dots$	Ho 156 7.8 m 9.5 s 56 m $\epsilon$ $\beta^+ \dots$ $\gamma 366;$ 266... $\gamma (52)$ $e^-$	Ho 157 12.6 m $\epsilon$ $\beta^+ 1.2; 1.5 \dots$ $\gamma 266; 193; 87 \dots$	Ho 158 21 m 27 m 11 m $\epsilon$ $\beta^+ \dots$ $\gamma 67; 99;$ 406; 218; 839; 946... $\gamma 206 \dots$	Ho 159 8.3 s 33 m $\epsilon; \beta^+ \dots$ $\gamma 121;$ 132; 253...	Ho 160 3 s 5.0 h 26 m $\epsilon; \beta^+ \dots$ $\gamma 118;$ 51; 107... $e^-$	Ho 161 6.7 s 2.5 h $\epsilon$ $\beta^+ \dots$ $\gamma 26;$ 78... $e^-$	Ho 162 68 m 15 m $\epsilon; \beta^+ \dots$ $\gamma 58; 38 \dots$ $e^-$	Ho 163 1.1 s 4570 a $\epsilon$ $\beta^+ 1.1 \dots$ $\gamma 81;$ 1220; 283; 937... $e^-$	Ho 164 37 m 29 m $\epsilon$ $\beta^+ 1.0 \dots$ $\gamma 91;$ 73... $e^-$	Ho 165 100 $\sigma 3.1 + 58$ $\sigma_n, \alpha < 2E-5$	Ho 166 1200 a 26.80 h $\beta^-$ 0.07... $\gamma 184;$ 810; 712 $\sigma 3100$ $e^-$
Dy 154 3.0 · 10 <sup>6</sup> a $\alpha 2.87$	Dy 155 10.0 h $\epsilon$ $\beta^+ 0.9; 1.1 \dots$ $\gamma 227 \dots$	Dy 156 0.056 $\sigma 33$ $\sigma_n, \alpha < 0.009$	Dy 157 8.1 h $\epsilon$ $\beta^+ \dots$ $\gamma 26 \dots$	Dy 158 0.095 $\sigma 33$ $\sigma_n, \alpha < 0.006$	Dy 159 44.4 d $\epsilon$ $\beta^+ \dots$ $\gamma 53; e^-$ $\sigma 3000$	Dy 160 2.329 $\sigma 60$ $\sigma_n, \alpha < 0.000$	Dy 161 18.889 $\sigma 600$ $\sigma_n, \alpha < 1E-5$	Dy 162 25.475 $\sigma 170$	Dy 163 24.896 $\sigma 120$ $\sigma_n, \alpha < 2E-5$	Dy 164 28.260 $\sigma 1610 + 1040$	Dy 165 1.3 m 2.35 h $\gamma 108; e^-$ $\beta^-$ 0.07... $\gamma 184;$ 810; 712 $\sigma 3100$ $e^-$
Tb 153 2.34 d $\epsilon$ $\beta^+ \dots$ $\gamma 212; 110; 102; 83 \dots$	Tb 154 23 h 9.0 h 21 h $\epsilon; \gamma$ $\gamma 248;$ 347; 123; 1420; 248; 123... $e^-$	Tb 155 5.32 d $\epsilon$ $\beta^+ \dots$ $\gamma 87; 105;$ 180; 262...	Tb 156 24 h? 5.4 h 5.4 d $\epsilon; \gamma$ $\gamma 88;$ $\gamma 534;$ 199; 1222 $e^-$	Tb 157 99 a $\epsilon$ $\beta^+ \dots$ $\gamma (54)$ $e^-$	Tb 158 10.5 s 180 a $\epsilon$ $\beta^- 0.9$ $\gamma 944;$ 962; 80... $e^-$	Tb 159 100 $\sigma 23.2$	Tb 160 723 d $\beta^- 0.6; 1.1 \dots$ $\gamma 879; 299;$ 966... $\sigma 570$	Tb 161 690 d $\beta^- 0.5; 0.6 \dots$ $\gamma 26; 49; 75 \dots$ $e^-$	Tb 162 7.76 m $\beta^- 1.4; 2.4 \dots$ $\gamma 260; 808;$ 888...	Tb 163 19.5 m $\beta^- 0.8; 1.3 \dots$ $\gamma 351; 390;$ 494...	Tb 164 3.0 m $\beta^- 1.7; 3.0 \dots$ $\gamma 169; 755;$ 215; 688; 611...
Gd 152 0.20 1.1 · 10 <sup>14</sup> a $\alpha 2.14; \sigma 700$ $\sigma_n, \alpha < 0.007$	Gd 153 239.47 d $\epsilon$ $\beta^+ 1.9; 1.1 \dots$ $\gamma 97; 103; 70 \dots$ $\sigma 20000$ $\sigma_n, \alpha 0.03$	Gd 154 2.18 $\sigma 60$	Gd 155 14.80 $\sigma 61000$ $\sigma_n, \alpha 0.000$	Gd 156 20.47 $\sigma \sim 2.0$	Gd 157 15.65 $\sigma 254000$ $\sigma_n, \alpha < 0.0$	Gd 158 24.84 $\sigma 2.3$	Gd 159 1748 h $\beta^- 1.0 \dots$ $\gamma 364; 58 \dots$	Gd 160 21.86 $\sigma 1.5$	Gd 161 3.6 m $\beta^- 1.6; 1.7 \dots$ $\gamma 361; 315;$ 102... $\sigma 20000$	Gd 162 8.2 m $\beta^- 1.0 \dots$ $\gamma 442; 403 \dots$	Gd 163 68 s $\beta^-$ $\gamma 288; 214;$ 1562; 1685...
Eu 151 47.81 $\sigma 4 + 3150 + 6000$	Eu 152 96 m 9.3 h 13.33 a $\beta^- 1.9 \dots$ $\epsilon; \beta^+ \dots$ $\beta^+ 0.7; 1.5$ $\gamma 841;$ 963; 344... $\sigma 68000$ $\sigma 11000$	Eu 153 52.19 $\sigma 300$ $\sigma_n, \alpha 1E-6$	Eu 154 46.0 m 8.8 a $\beta^- 0.6; 1.8 \dots$ $\epsilon; \gamma 123$ 1274; 723; 1005... $\sigma 1500$	Eu 155 4761 a $\beta^- 0.17; 0.5 \dots$ $\gamma 87; 105 \dots$ $\sigma 3900$	Eu 156 5.2 d $\beta^- 0.5; 2 \dots$ $\gamma 812; 89;$ 1231...	Eu 157 5.18 h $\beta^- 1.3 \dots$ $\gamma 64; 411;$ 371; 619...	Eu 158 46 m $\beta^- 2.4; 3 \dots$ $\gamma 944; 977; 898$	Eu 159 1.1 m $\beta^- 2.6 \dots$ $\gamma 68; 71; 79;$ 96; 103...	Eu 160 12 s $\beta^- 4.1 \dots$ $\gamma 173; 515;$ 412; 822...	Eu 161 26 s $\beta^-$ $\gamma 72-314$	Eu 162 10.6 s $\beta^-$ $\gamma 71; 165$
Sm 150 7.38 $\sigma 102$	Sm 151 93 a $\beta^- 0.1 \dots$ $\gamma (22 \dots); e^-$ $\sigma 15200$	Sm 152 26.75 $\sigma 206$	Sm 153 46.27 h $\beta^- 0.7; 0.8 \dots$ $\gamma 103; 70 \dots$ $\sigma 420$	Sm 154 22.75 $\sigma 7.5$	Sm 155 2.4 m $\beta^- 1.5 \dots$ $\gamma 104; 246;$ 141...	Sm 156 1.5 m $\beta^- 0.7 \dots$ $\gamma 204; 88; 166 \dots$ $e^-$	Sm 157 1.5 m $\beta^- 2.4 \dots$ $\gamma 198; 196;$ 394...	Sm 158 1.5 m $\beta^- 1.89; 364;$ 325...	Sm 159 1.5 m $\beta^- 1.56; 302;$ 254; 797; 179...	Sm 160 1.5 m $\beta^-$ $\gamma 110 \dots$	Sm 161 4.8 s $\beta^-$ $\gamma 264$

s and r process

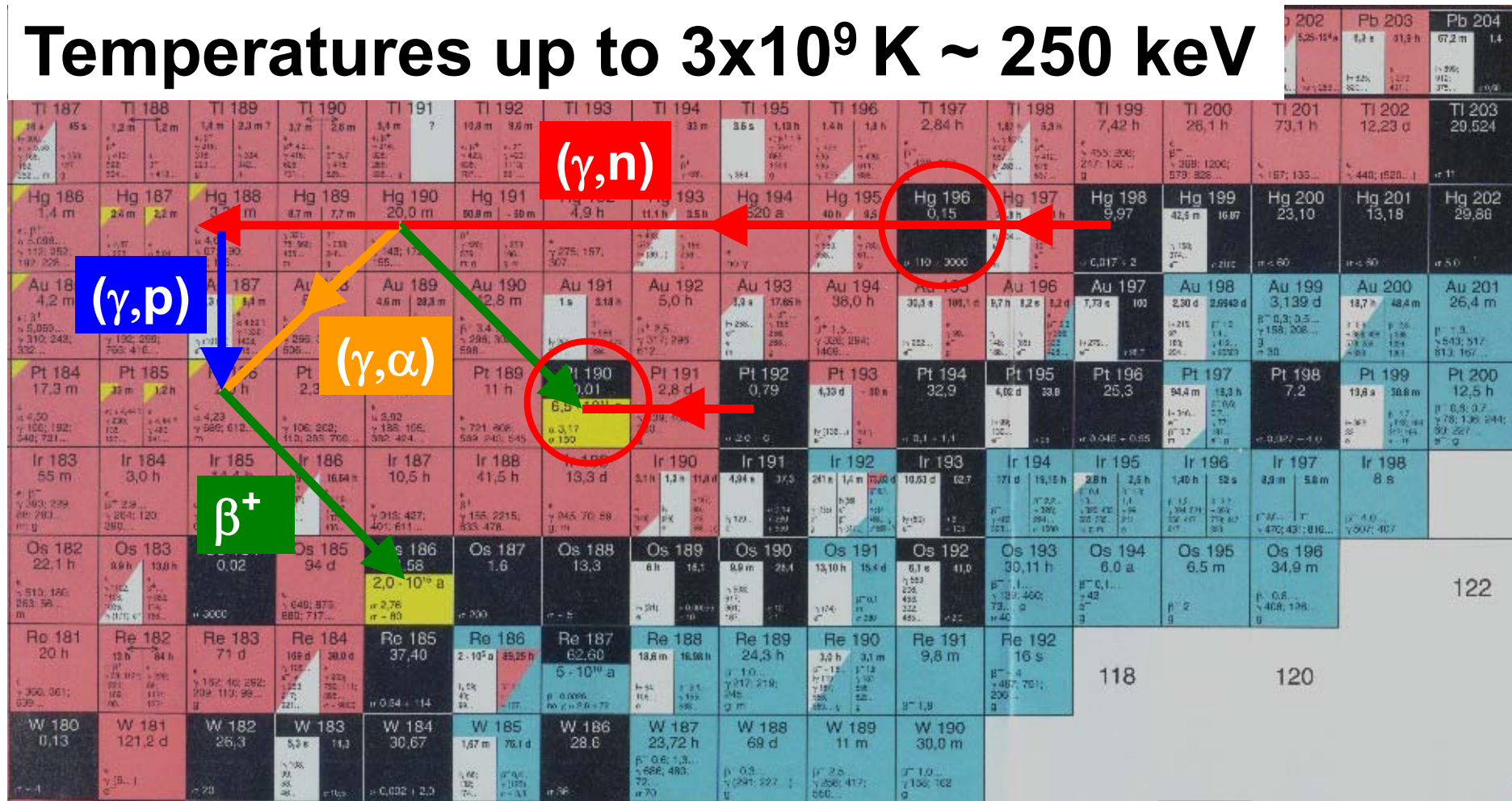
# *p* nuclei

- 35 proton rich stable nuclei with  $A > 70$
- typical isotopic abundance  $< 1\%$
- typical abundance  $5 \times 10^{-4}$  ( $\text{Si} = 10^6$ )



# The $p$ process: Nuclear reaction network

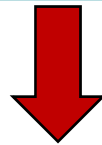
Temperatures up to  $3 \times 10^9$  K  $\sim$  250 keV



For lighter nuclei there may be competing reactions:  
 $(n, \gamma)$ ,  $(p, \gamma)$ ,  $(\alpha, \gamma)$ ,  $\nu p$ -process

# The $p$ process: Nuclear reaction network

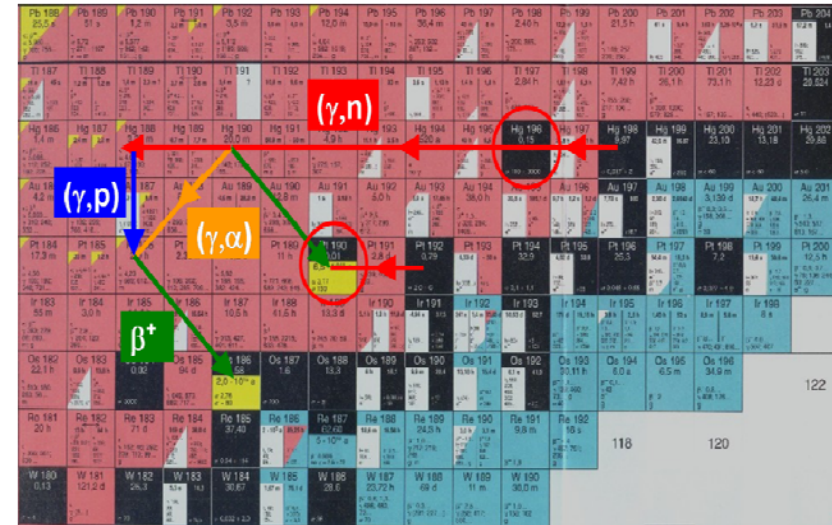
This network involves about 20.000 reaction rates



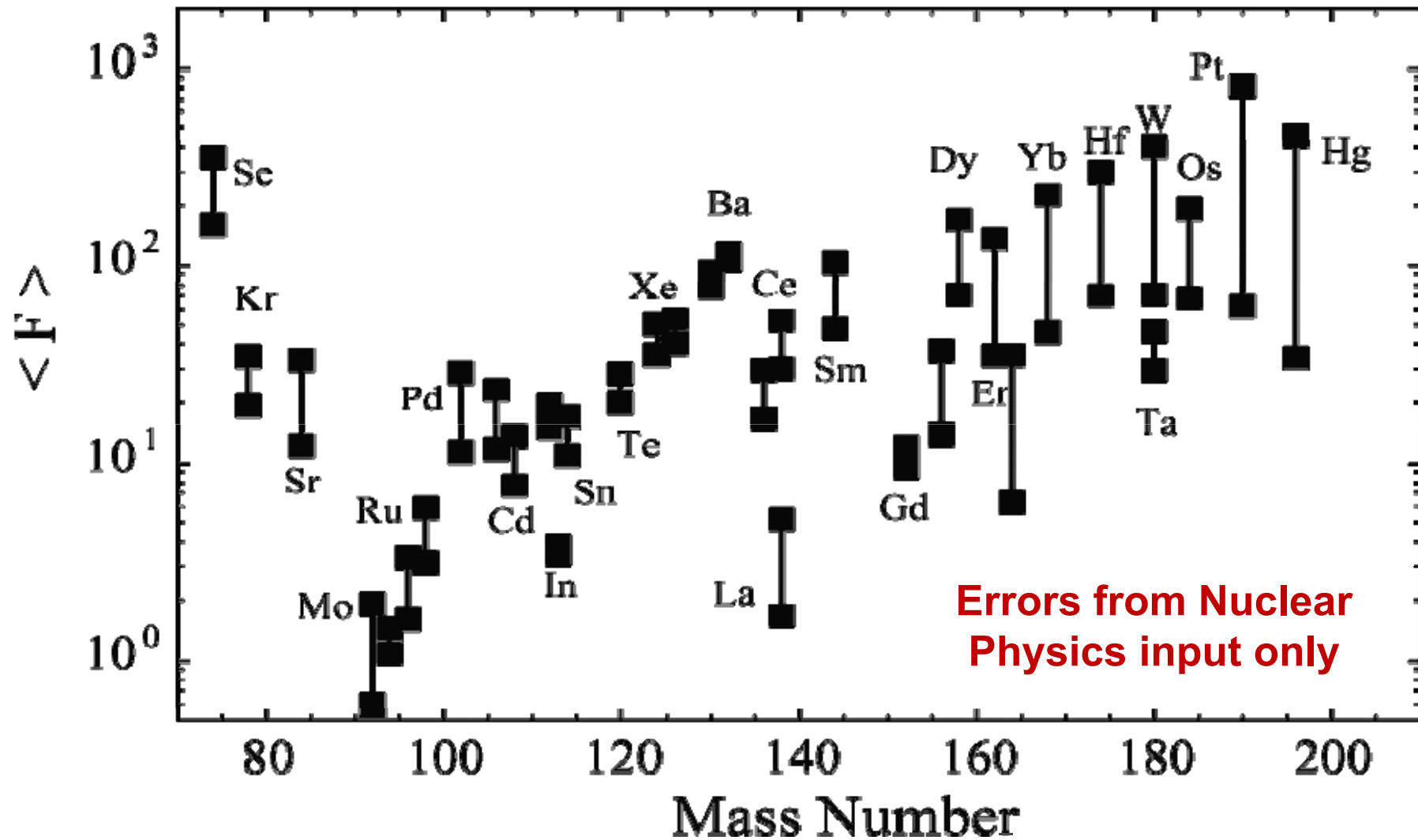
Statistical model calculations



Experiments are necessary to check and to constrain the model assumptions



# Abundance of $p$ nuclei: Prediction vs. observation



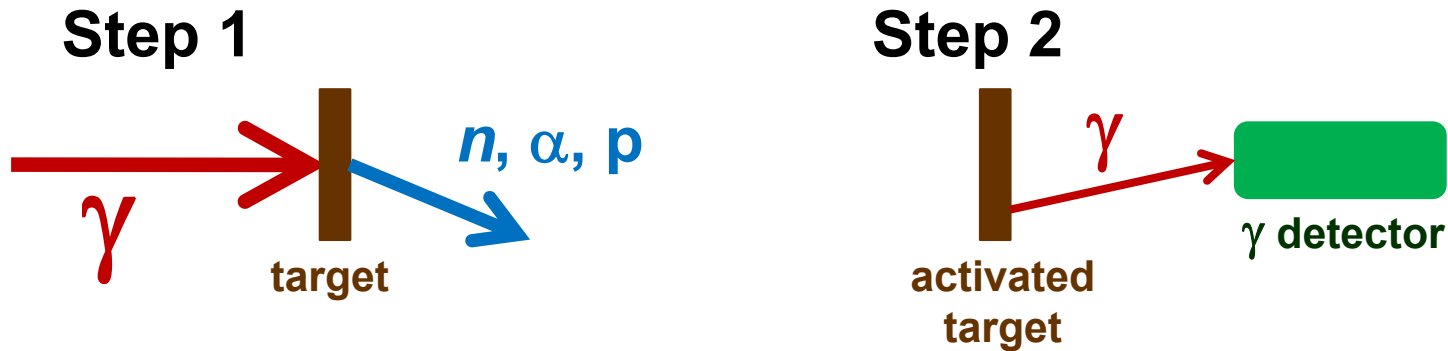
*M. Arnould and S. Goriely, Phys. Rep. 384 (2003) 1*

*S. Goriely et al., Astronomy & Astrophysics 444 (2005) L1*

# Nuclear Physics input for the $p$ process

- Ground state masses
- Properties of excited states
- Level densities
- Photoresponse  $(\gamma, \gamma')$ ,  $(\gamma, n)$ ,  $(\gamma, \alpha)$ ,  $(\gamma, p)$
- Optical potentials (e.g.  $\alpha$  – nucleus)

# „Direct“ experimental techniques I: Activation of stable targets with real photons

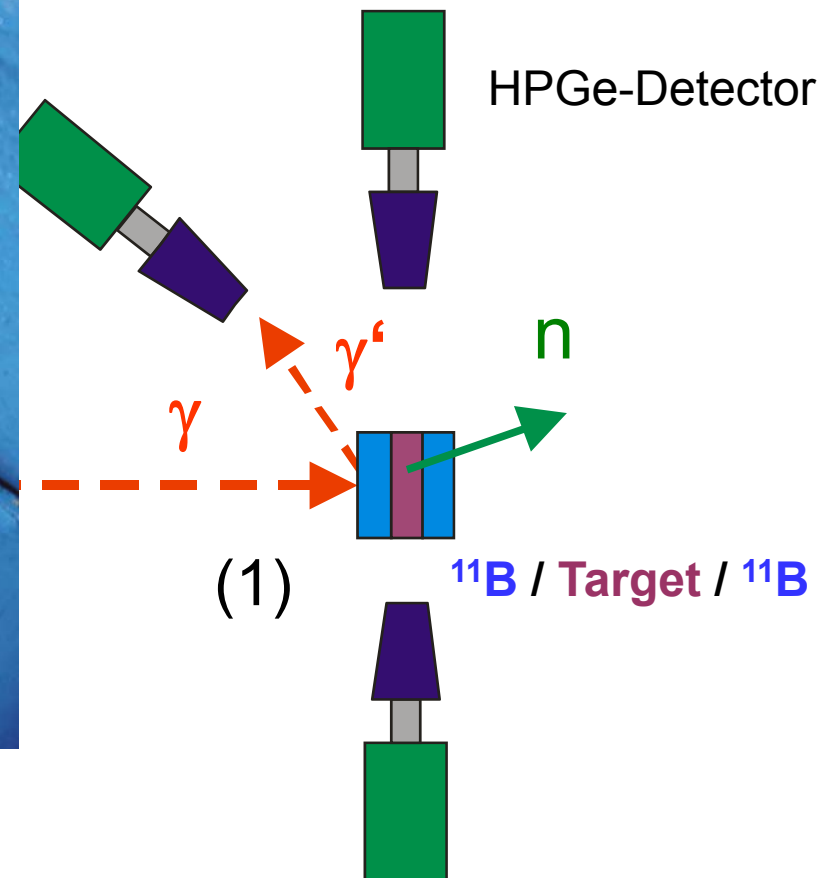
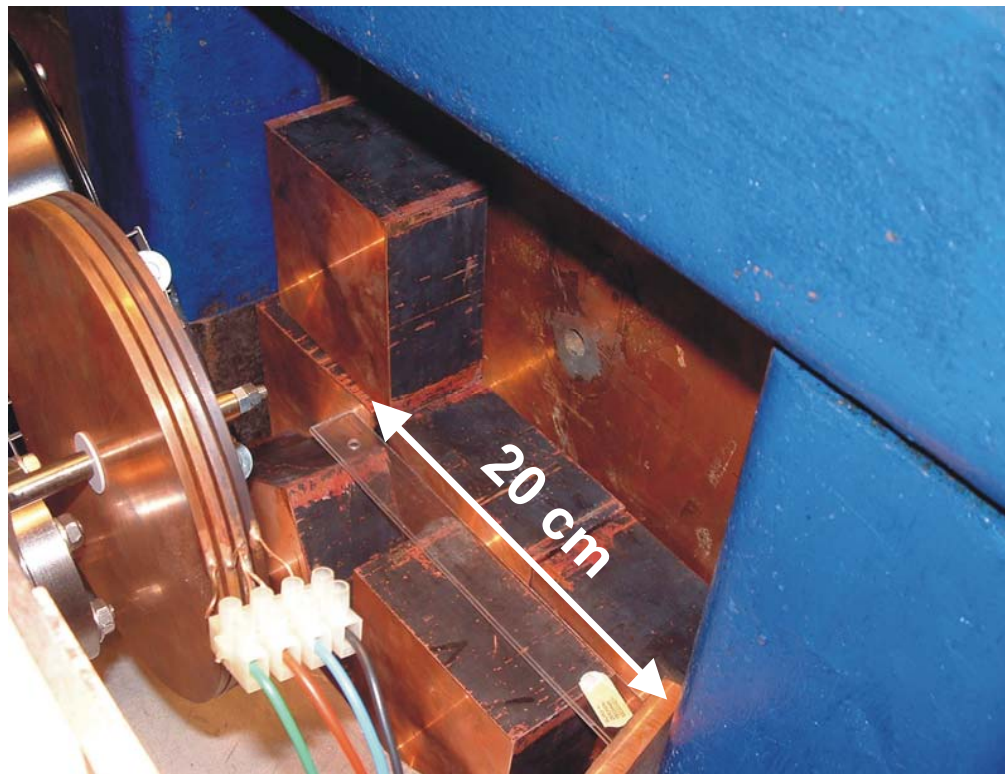


Induce  $(\gamma, n)$ ,  $(\gamma, \alpha)$ , and  $(\gamma, p)$  reactions with real photons + measure activation

Er 161 3.24 h $\epsilon$ $\beta^+$ ... $\gamma$ 827... g; m	Er 162 0.139 $\sigma$ 19 $\sigma_n, \alpha < 0.011$	Er 163 75 m $\epsilon$ $\beta^+$ ... $\gamma$ (1114...) g	Er 164 1.601 $\sigma$ 13 $\sigma_n, \alpha < 0.0012$	Er 165 10.3 h $\epsilon$ no $\gamma$	Er 166 33.503 $\sigma$ 3 + 14 $\sigma_n, \alpha < 7E-5$
Ho 160 3 s   5.0 h   26 m $\epsilon$ $\beta^+$ ... $\gamma$ 728; 879; 962... ly 118; 51; 107... $e^-$ 2974	Ho 161 6.7 s   2.5 h $\epsilon$ $\gamma$ 26; 78... $e^-$	Ho 162 68 m   15 m $\epsilon$ $\beta^+$ 1.1... $\gamma$ 185; 1220; 283; 1319... 937... $e^-$	Ho 163 1.1 s   4570 a $\epsilon$ no $\gamma$	Ho 164 37 m   29 m $\epsilon$ $\beta^-$ 1.0... $\gamma$ 91; 73... $e^-$	Ho 165 100 $\sigma$ 3.1 + 58 $\sigma_n, \alpha < 2E-5$
Dy 159 144.4 d $\epsilon$ $\gamma$ 58; $e^-$ $\sigma$ 8000	Dy 160 2.329 $\sigma$ 60 $\sigma_n, \alpha < 0.0003$	Dy 161 18.889 $\sigma$ 600 $\sigma_n, \alpha < 1E-6$	Dy 162 25.475 $\sigma$ 170	Dy 163 24.896 $\sigma$ 120 $\sigma_n, \alpha < 2E-5$	Dy 164 28.260 $\sigma$ 1610 + 1040



# Photoactivation with bremsstrahlung

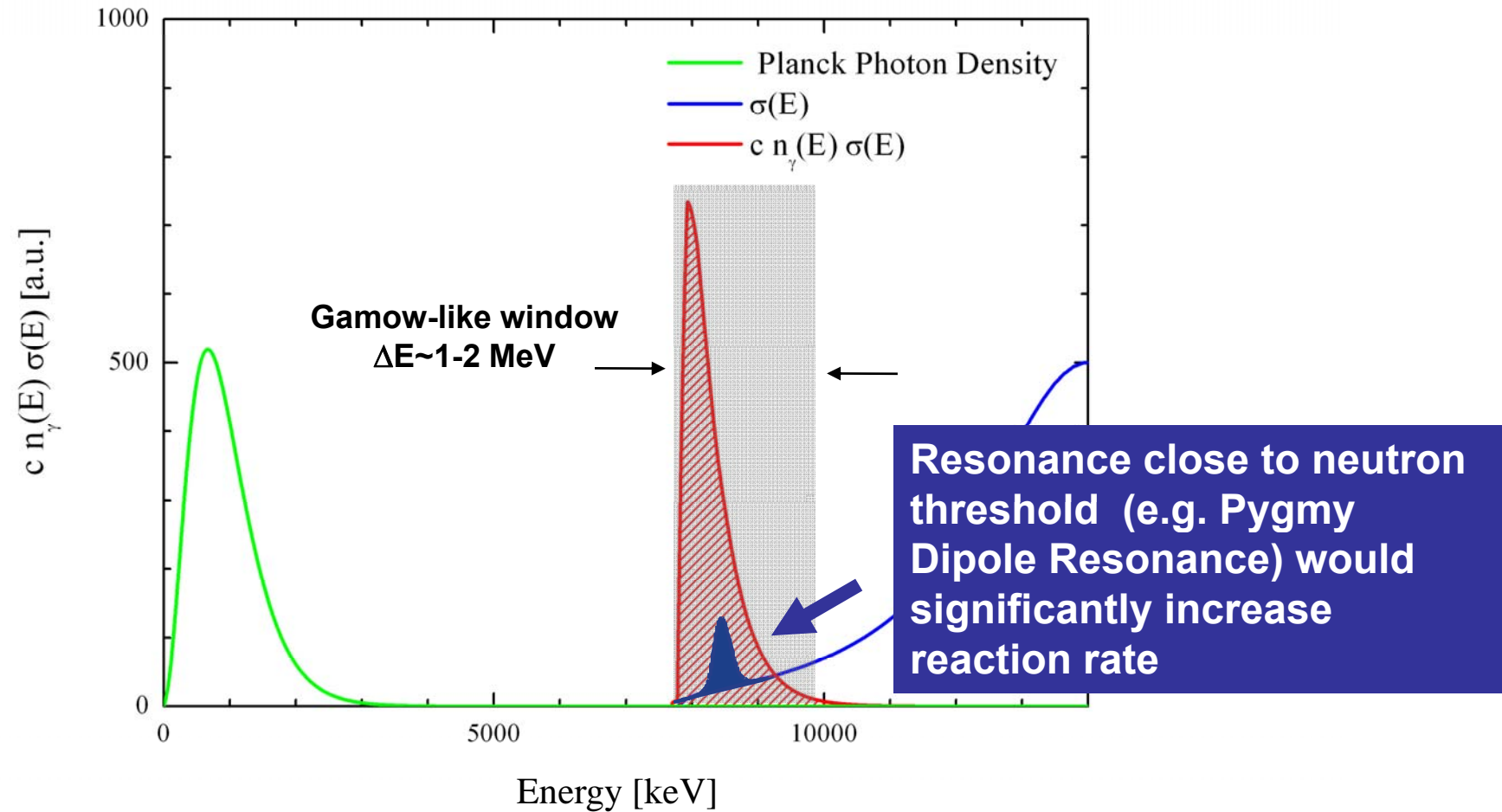


(1) Photon flux  $\sim 10^5 \gamma / (\text{keV s cm}^2)$   
Calibration of the photon flux via  $^{11}\text{B}(\gamma, \gamma')$

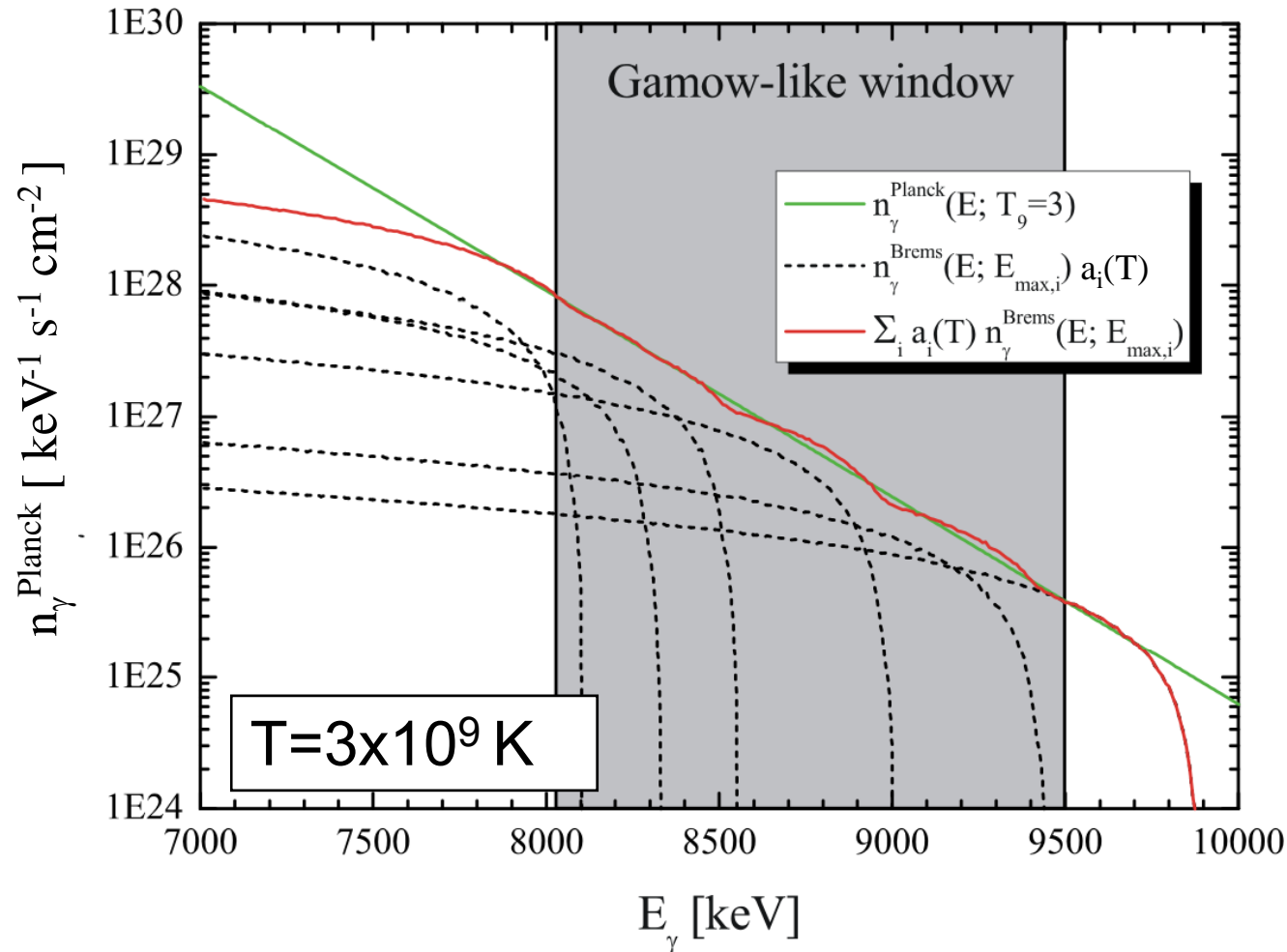
(2) Photon flux  $\sim 10^7 \gamma / (\text{keV s cm}^2)$   
Calibration of the photon flux via  $^{197}\text{Au}(\gamma, n)$  and  $^{187}\text{Re}(\gamma, n)$

# Energy region of interest: Gamow window for $(\gamma, n)$

Reaction rate:  $\lambda(T) = c \int n_\gamma(E) \sigma(E) dE$



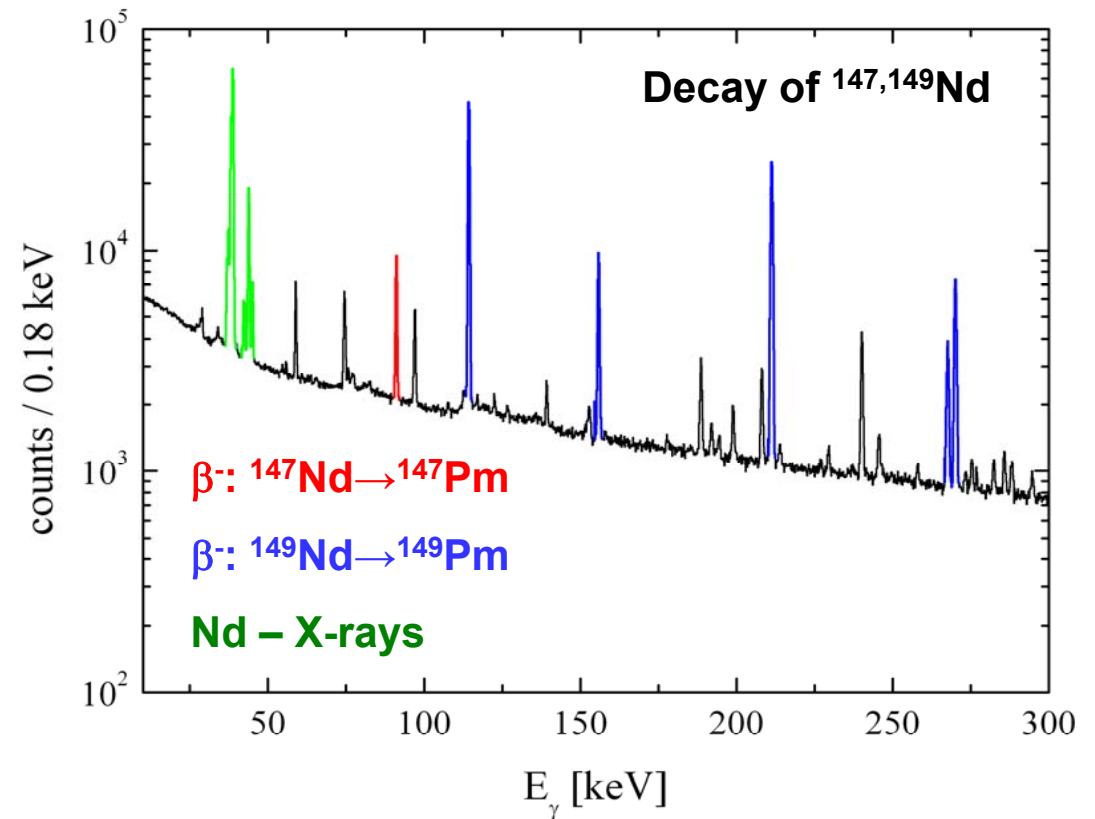
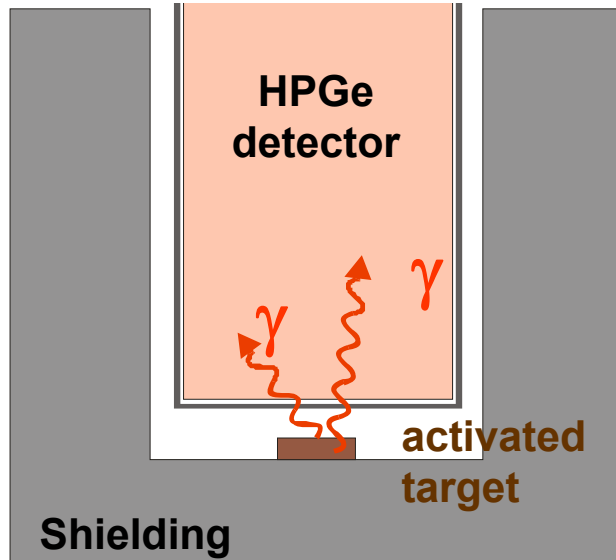
# Production of a quasi-thermal spectrum



A. Z. *et al.*, Prog. Part. Nucl. Phys. 44 (2000) 39

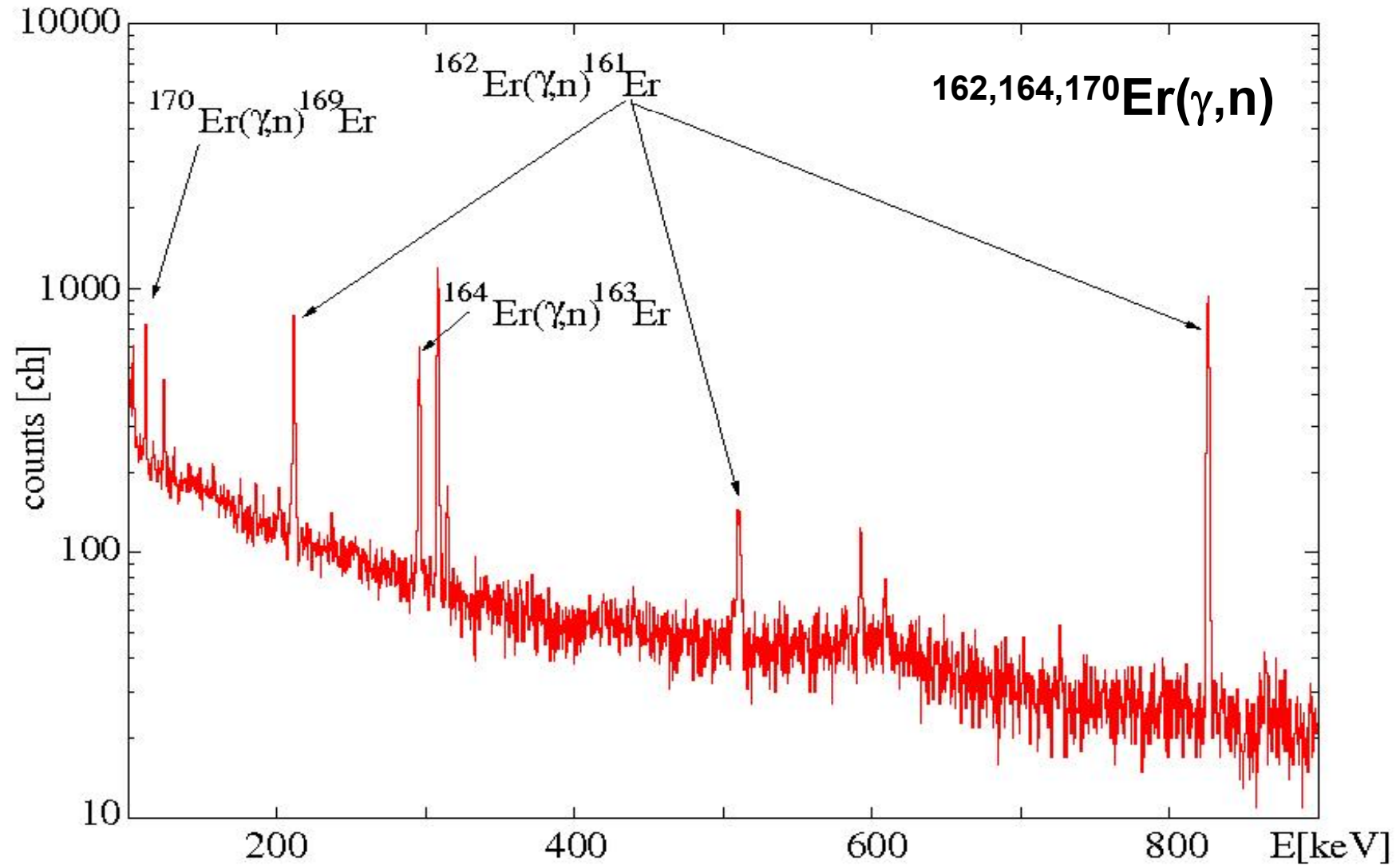
P. Mohr *et al.*, Phys. Lett. B **488** (2000) 127

# Activation measurement



Reaction Yield:  $Y \sim \int n_\gamma(E) \sigma(E) dE$

# Photodissociation of Er isotopes



*J. Hasper and S. Müller, priv. comm.*

# Groundstate reaction rates @ $2.5 \times 10^9$ K

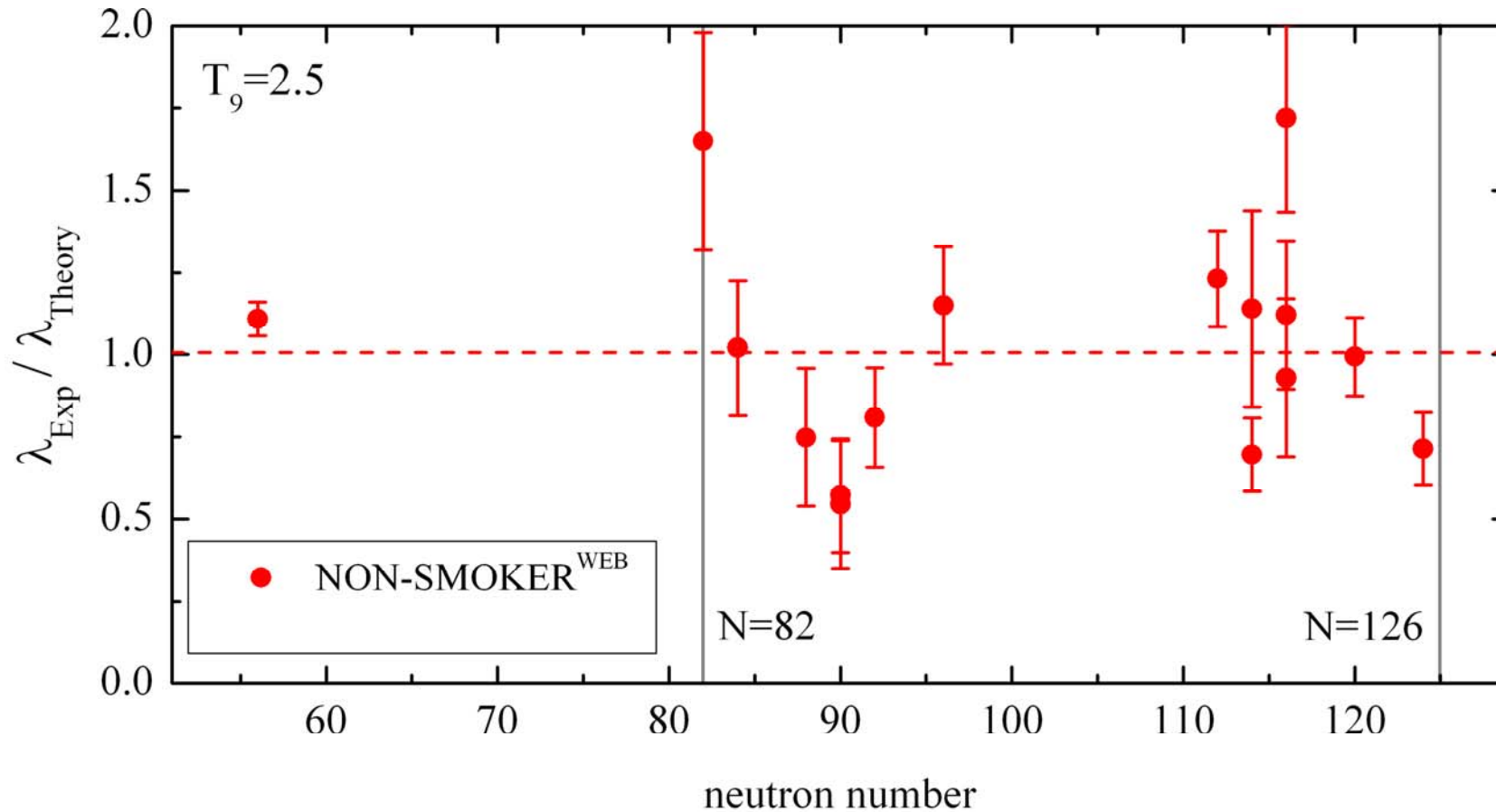
Isotope	$\lambda_{\text{exp,gs}}$	Reference	$\lambda_{\text{NONS,gs}}$	$\lambda_{\text{TALYS,gs}}$
$^{148}\text{Nd}$	65(18)	J. Hasper et al., PRC 77 (2008) 015803	86	82
$^{150}\text{Nd}$	41(12)		72	77
$^{154}\text{Sm}$	7.5(1.4)		9.2	11
$^{187}\text{Re}$	91(11)	S. Müller et al., PRC 73 (2006) 025804	74	99
$^{190}\text{Pt}$	0.6(2)	K. Vogt et al., PRC 63 (2001) 055802	0.14	0.20
$^{192}\text{Pt}$	0.4(1)		0.53	0.62
$^{198}\text{Pt}$	61(7)		61	80
$^{197}\text{Au}$	6.2(8)	K. Vogt et al., NPA 707 (2002) 241	5.3	7.5
$^{196}\text{Hg}$	0.43(7)	K. Sonnabend et al., PRC 70 (2004) 035802	0.25	0.36
$^{198}\text{Hg}$	2.0(3)		1.41	1.75
$^{204}\text{Hg}$	57(9)		80.1	108
$^{191}\text{Ir}$	4.3(5)	J. Hasper, submitted to PRC	4.5	4.6
$^{193}\text{Ir}$	13.5(16)		16	24

**NON-SMOKER:**  
T. Rauscher and  
F.-K. Thielemann,  
ADNDT 75 (2000) 1

**TALYS:**  
A. Koning et al.

(all reaction rates in  $\text{s}^{-1}$ )

# Groundstate reaction rates @ $2.5 \times 10^9 \text{K}$



**NON-SMOKER<sup>WEB</sup> and TALYS deviate by about 20%**

# From integrated reaction rates to $\sigma(E)$

Untagged photons from bremsstrahlung  
measure always INTEGRATED reaction rates:

$$\lambda(T) = c \int n_{\gamma}(E) \sigma(E) dE$$

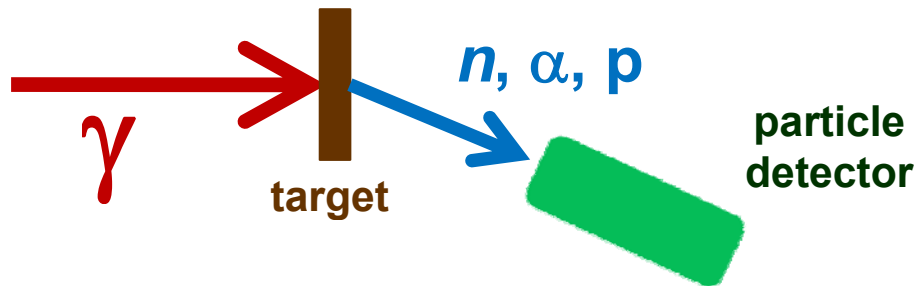
Additional information can be deduced from  
the shape of the cross section  $\sigma(E_{\gamma})$

→ use photons with „known“ energy

- Tagged bremsstrahlung photons
- Laser Compton Backscattering

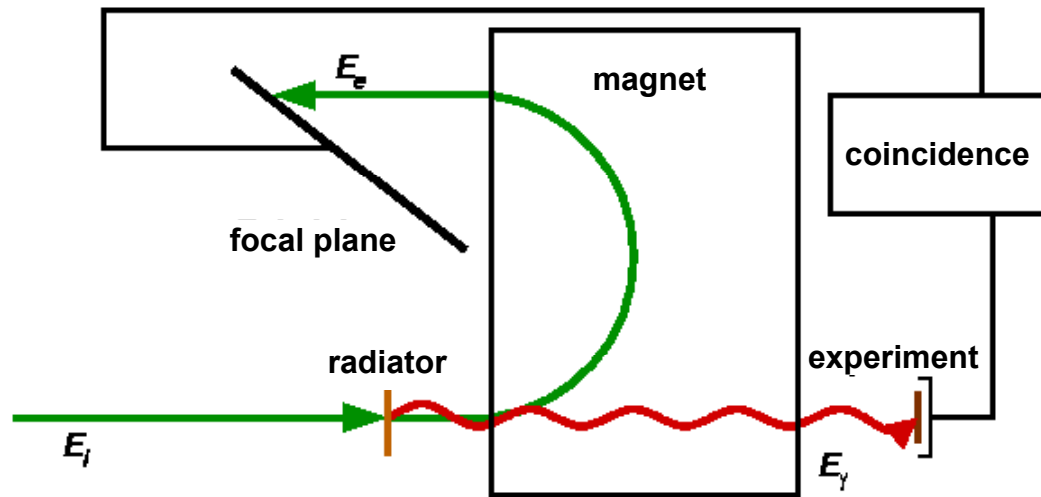


# „Direct“ experimental techniques II: Dissociation with monoenergetic photons



**Induce  $(\gamma, n)$ ,  $(\gamma, \alpha)$ , and  $(\gamma, p)$  reactions with  
real photons with known energy +  
measure  $n$ ,  $\alpha$ , or  $p$**

# Photon tagger NEPTUN @ S-DALINAC

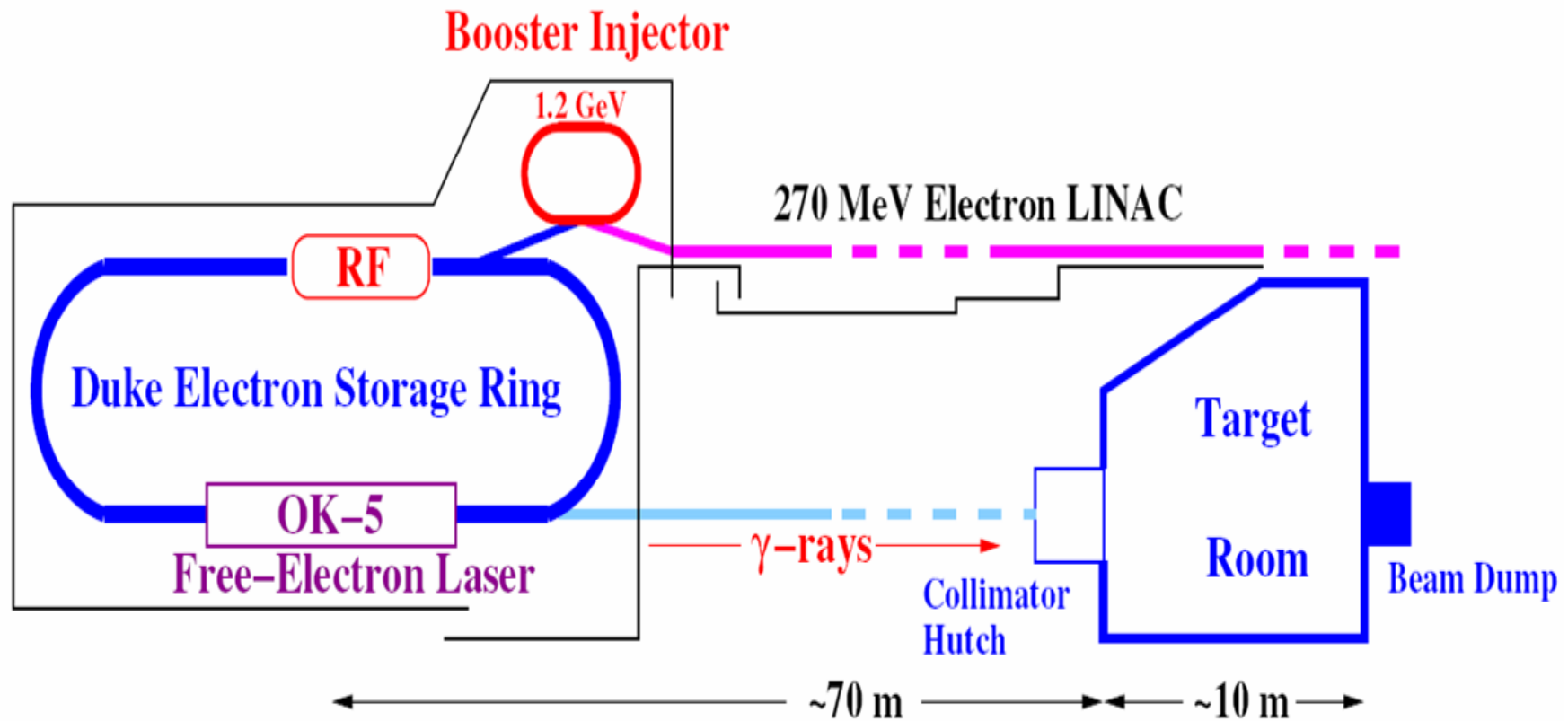


- $6 \text{ MeV} \leq E_\gamma \leq 20 \text{ MeV}$
- $\Delta E = 25 \text{ keV} @ 10 \text{ MeV}$
- Photon intensity:  $\approx 10^4 \text{ keV}^{-1}\text{s}^{-1}$

→ Measure  $(\gamma, \gamma')$ ,  $(\gamma, n)$ ,  $(\gamma, p)$ ,  
and  $(\gamma, \alpha)$  cross sections



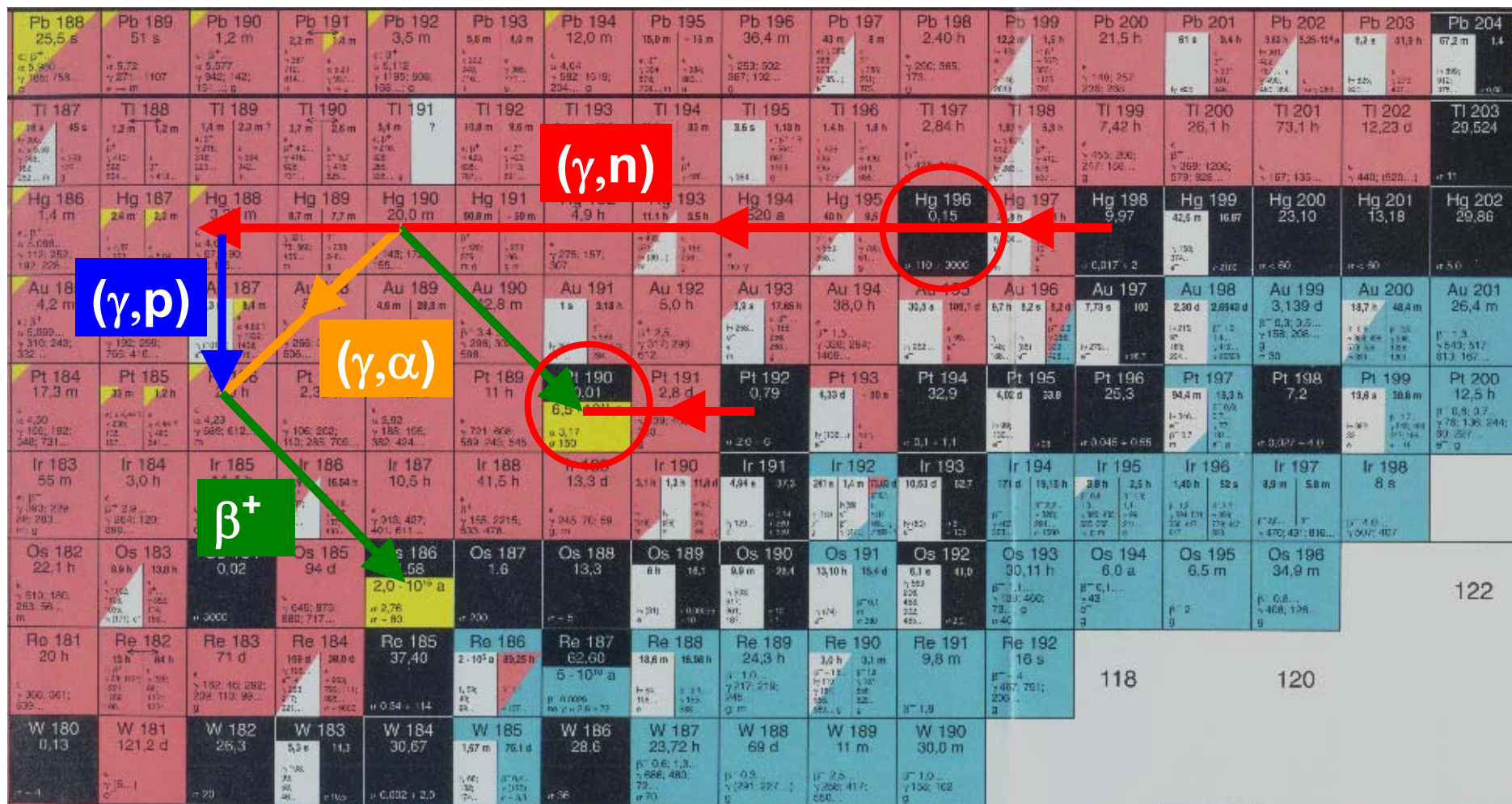
# Laser Compton Backscattering at H $\gamma$ S (Duke)



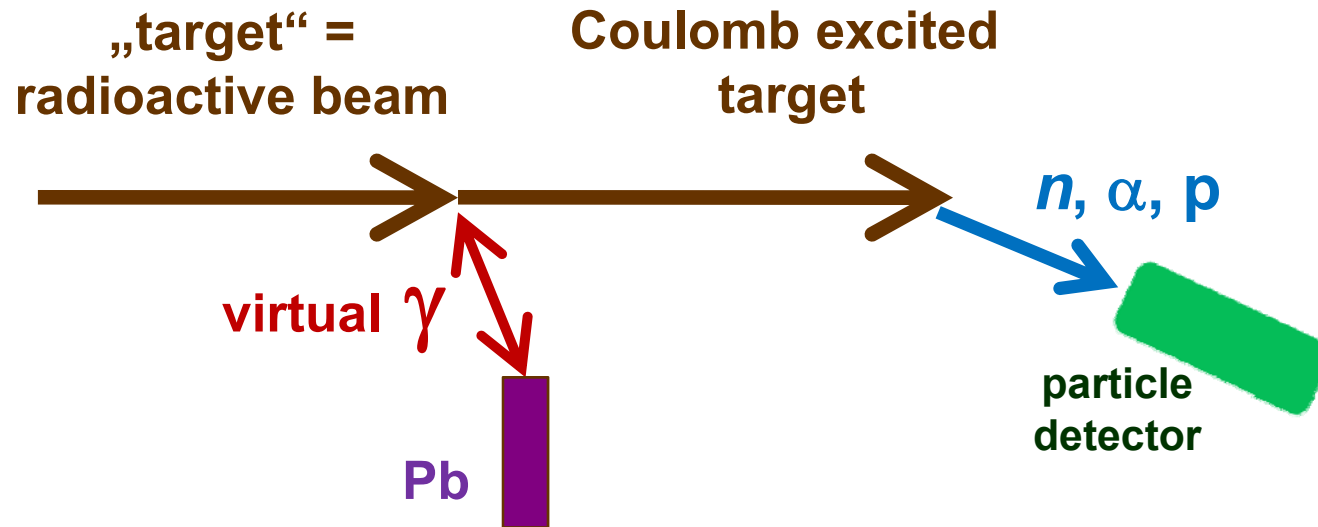
→ see next talk by Anne Sauerwein

# The limitations of real photon experiments

The above methods are limited to „stable“ target nuclei, but:

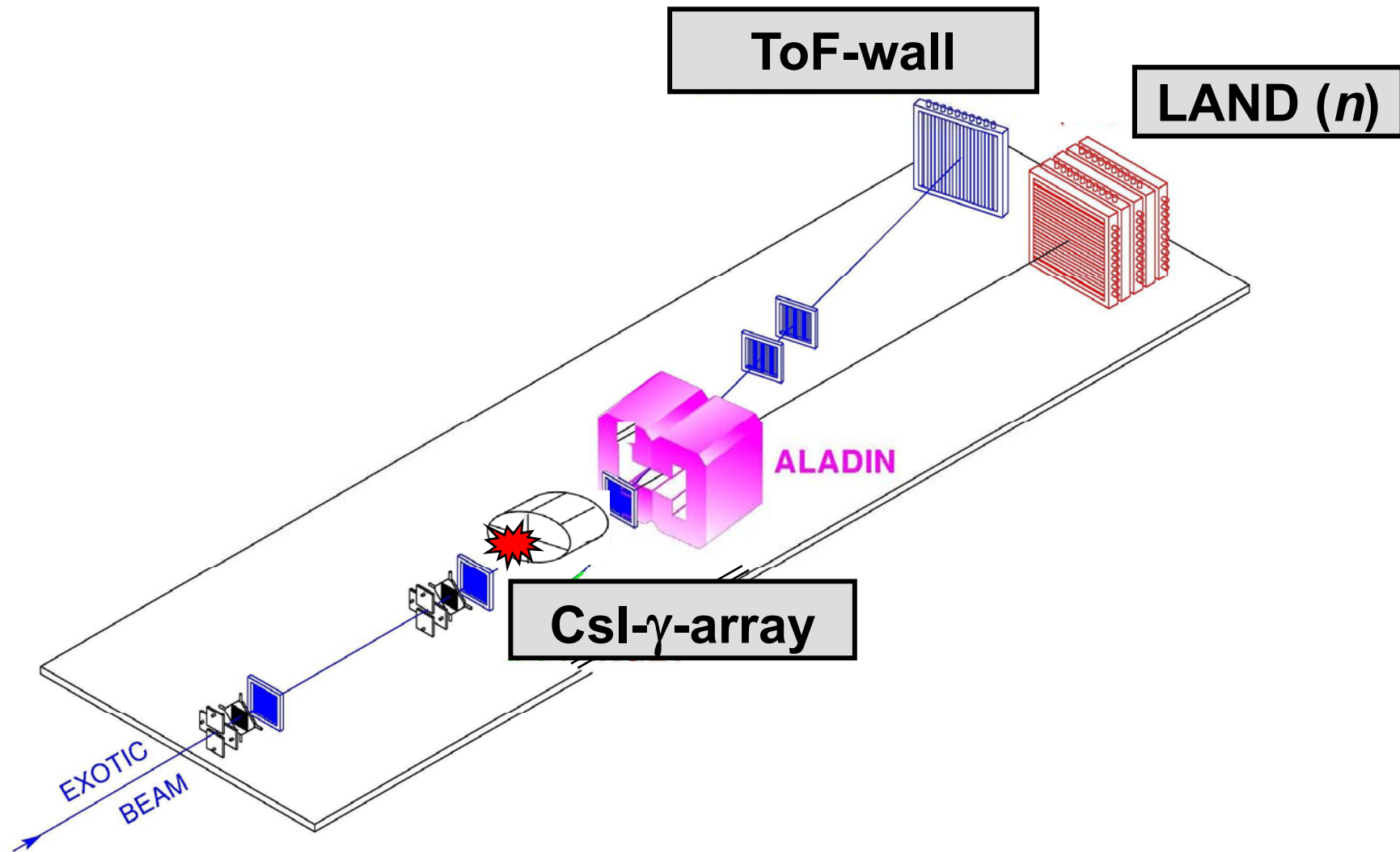


# „Direct“ experimental techniques III: Coulomb dissociation



Induce  $(\gamma, n)$ ,  $(\gamma, \alpha)$ , and  $(\gamma, p)$  reactions with virtual photons („inverse kinematics“)  
+ measure footprints of reaction

# (Super)FRS / (Neu)LAND setup at GSI / FAIR



*T. Aumann, Eur. Phys. Journal A 26 (2005) 441*

# Summary: Measurements with photons

## Activation experiments with bremsstrahlung

- Planck-like photon spectrum
- Integrated cross sections are determined
- Daughter nucleus must show observable decay
- Limited to „stable“ nuclei

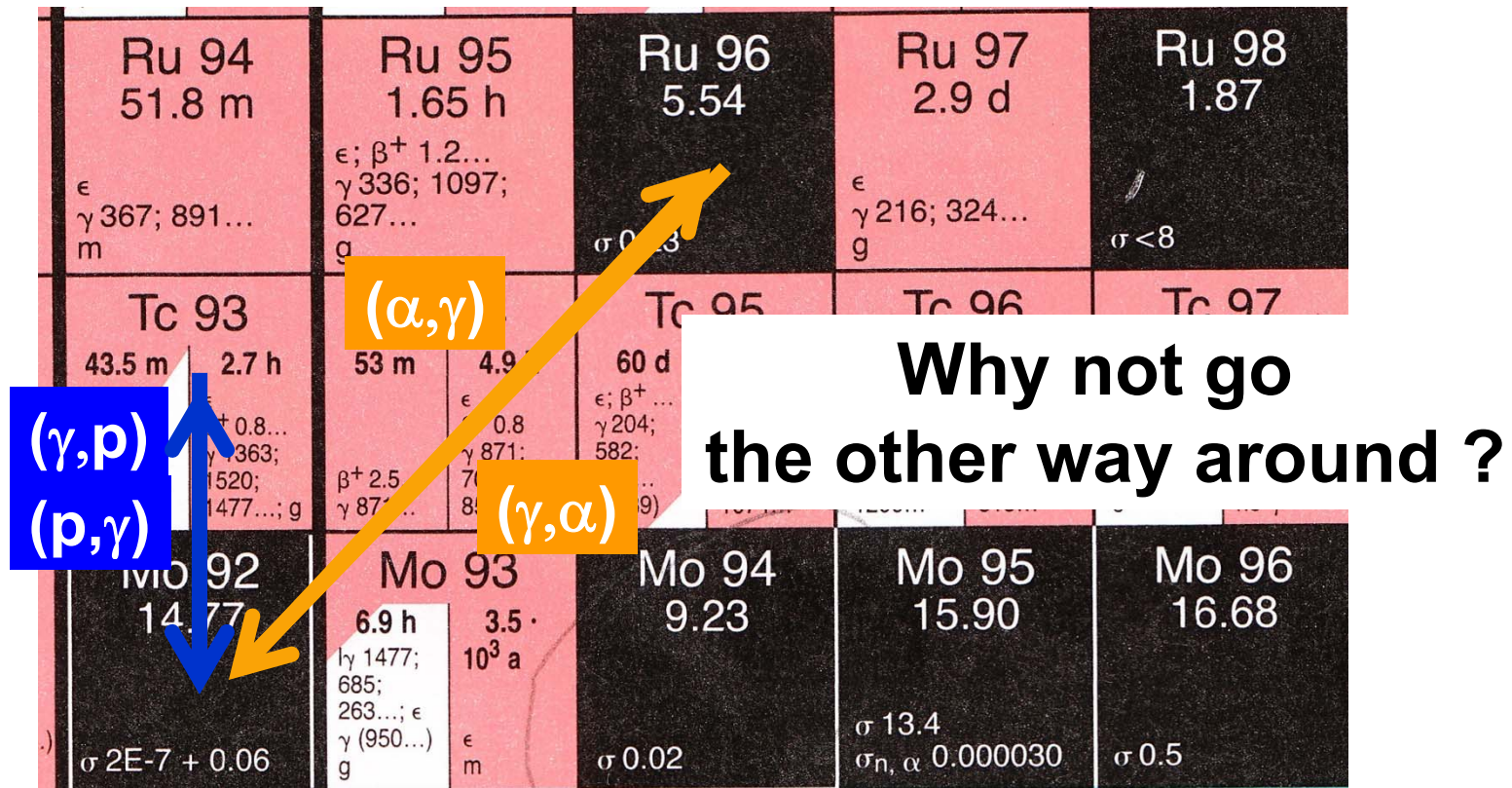
## Experiments with monoenergetic photons (Tagger or laser Compton backscattering)

- Energy dependent cross section  $\sigma(E)$  is observed
- Difficult experiments (photon flux !)
- Limited to „stable“ nuclei

# Summary: Measurements with photons

## Coulomb dissociation with radioactive beams

- Energy dependent cross section is observed
- Very difficult and complex experiments





# Indirect experimental methods: Radiative capture

Measurements of inverse capture reactions like  $(p,\gamma)$  and  $(\alpha,\gamma)$  can:

- yield additional information (detailed balance)
- determine the  $p$ -nucleus or  $\alpha$ -nucleus potential
- allow to study radioactive nuclei

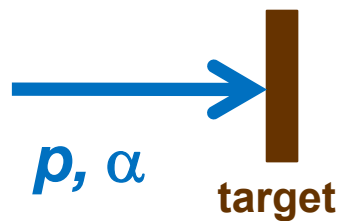
Ru 94 51.8 m $\epsilon$ $\gamma$ 367; 891... m		Ru 95 1.65 h $\epsilon$ ; $\beta^+$ 1.2... $\gamma$ 336; 1097; 627... g		Ru 96 5.54 $\sigma$ 0.23		Ru 97 2.9 d $\epsilon$ $\gamma$ 216; 324... g		Ru 98 1.87 $\sigma < 8$	
Tc 93 43.5 m   2.7 h $\beta^+$ 0.8... $\gamma$ 1363; 1520; 1477...; g		Tc 94 53 m   4.9 h $\epsilon$ $\beta^+$ 0.8 $\gamma$ 871; 703; 850... $\beta^+$ 2.5... $\gamma$ 871...		Tc 95 60 d   20 h $\epsilon$ ; $\beta^+$ ... $\gamma$ 204; 582; 835... $\beta^+$ 0.8 $\gamma$ 766; 1074... $\beta^+$ 2.5... $\gamma$ 871...		Tc 96 52 m   4.3 d $\beta^+$ 0.8 $\gamma$ 778; 850; 813... $\beta^+$ 0.8 $\gamma$ 778; 850; 813... $\beta^+$ 0.8 $\gamma$ 778; 850; 813...		Tc 97 92.2 d   4.0 · 10 <sup>6</sup> a $\beta^+$ 0.8 $\gamma$ 778; 850; 813... $\beta^+$ 0.8 $\gamma$ 778; 850; 813...	
Mo 92 14.77 $\sigma$ 2E-7 + 0.06		Mo 93 6.9 h   3.5 · 10 <sup>3</sup> a $\beta^+$ 0.8 $\gamma$ 1477; 685; 263...; $\epsilon$ $\gamma$ (950...) g		Mo 94 9.23 $\sigma$ 0.02		Mo 95 15.90 $\sigma$ 13.4 $\sigma_n, \alpha$ 0.000030		Mo 96 16.68 $\sigma$ 0.5	

# Indirect experimental methods I: Activation experiments after radiative capture

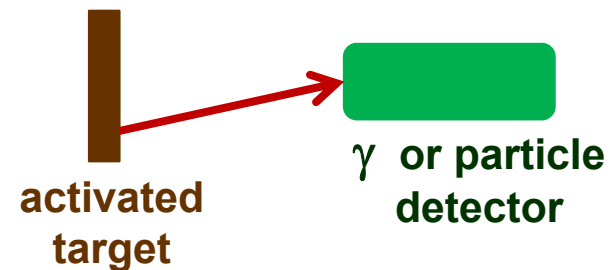
## Activation experiments

(e.g., G. G. Kiss et al., *PRC* 76 (2007) 055807)

Step 1



Step 2



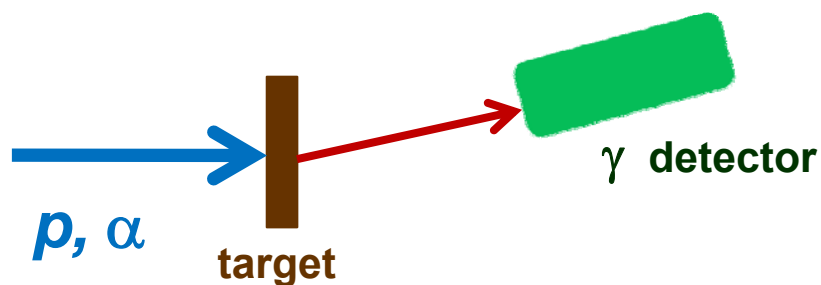
**Very sensitive but not always possible.**

# Indirect experimental methods II: In-beam measurements after radiative capture

## In-beam $\gamma$ spectroscopy

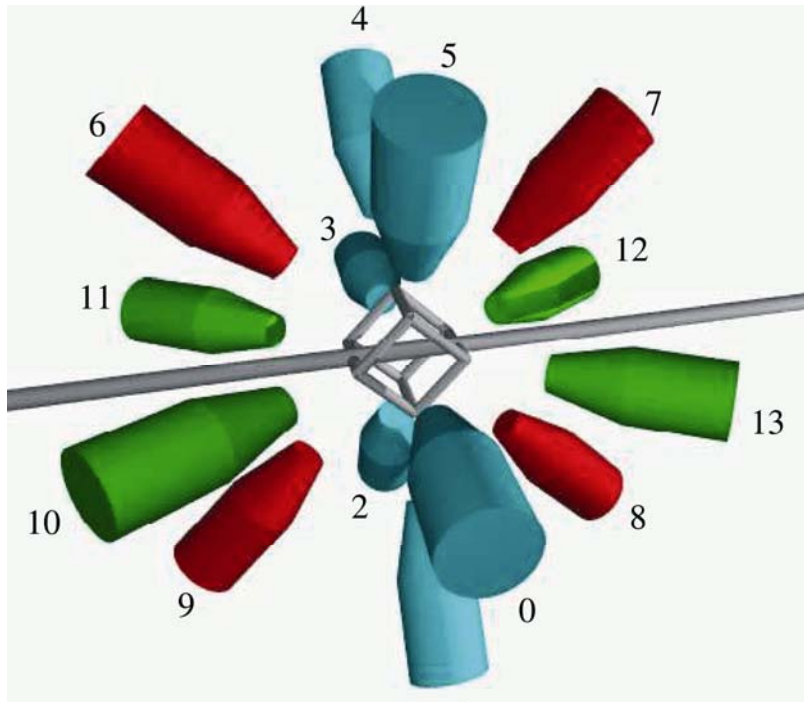
(e.g. with  $4\pi$  NaI: A. Spyrou et al., PRC 76 (2007) 015802)

(e.g. with HPGe: S. Ganalopoulos et al., PRC 67 (2003) 015801)



**Less sensitive but very versatile.**

# The HORUS $\gamma$ -detector array @ IKP Köln

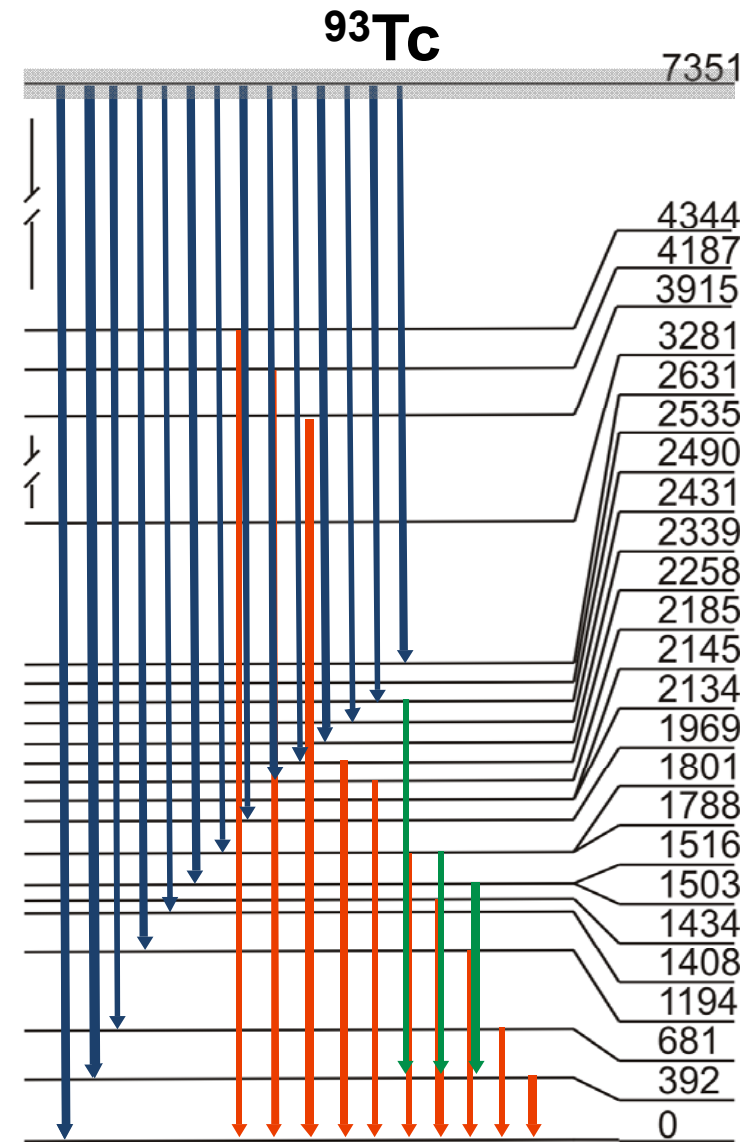
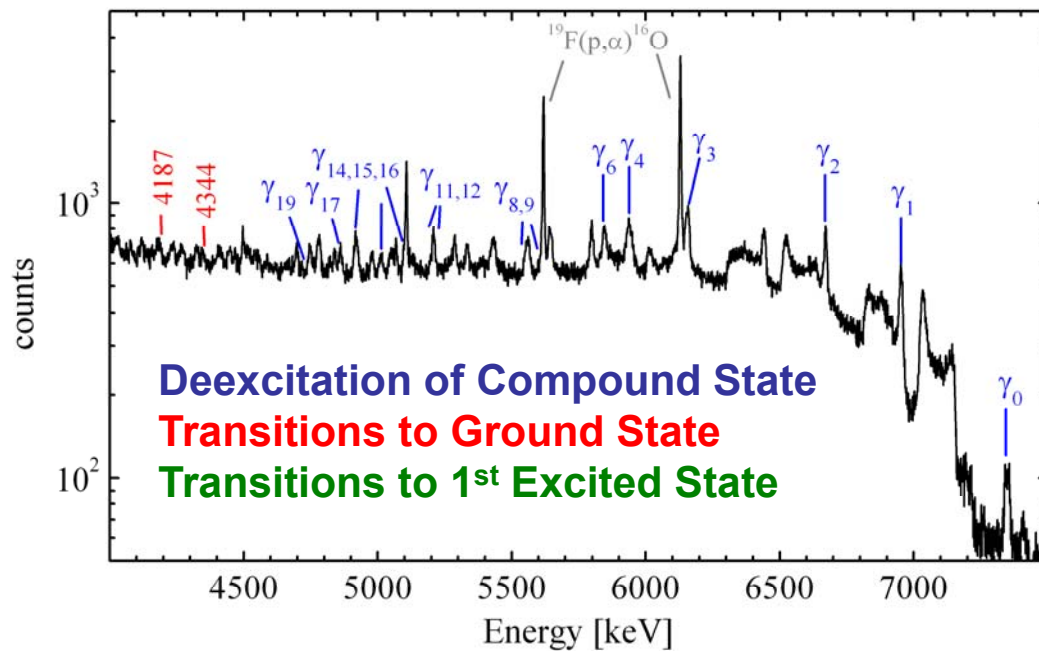
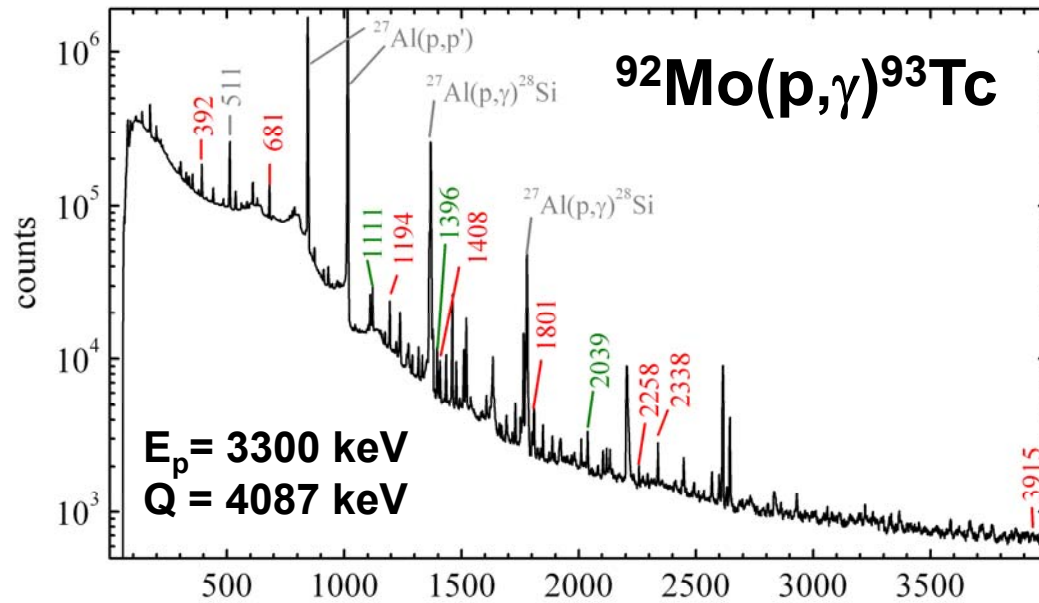


**14 HPGe detectors in close geometry**

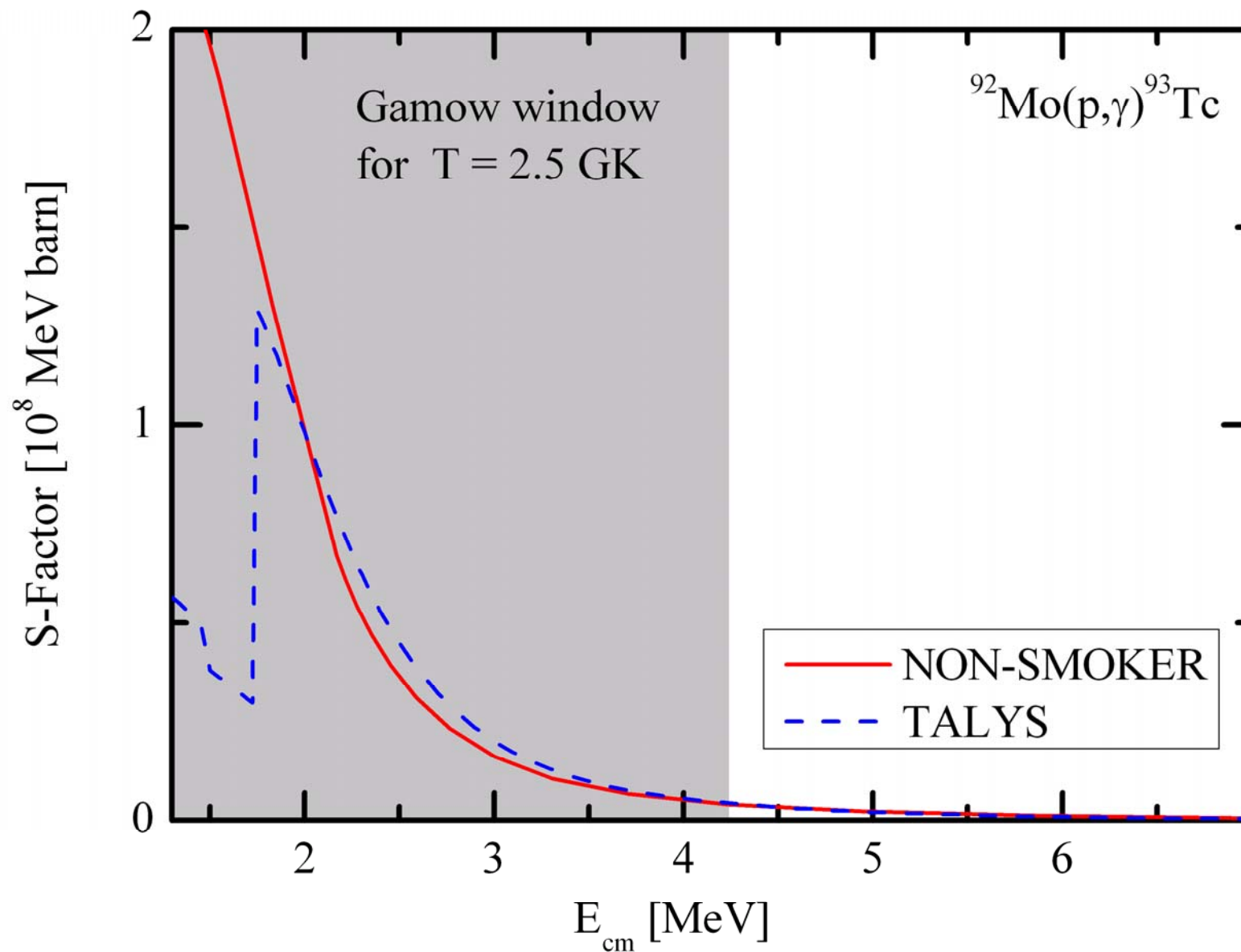
⇒ **Photopeak efficiency at 1332 keV: ~ 2%**

⇒ **No requirements on reaction products**

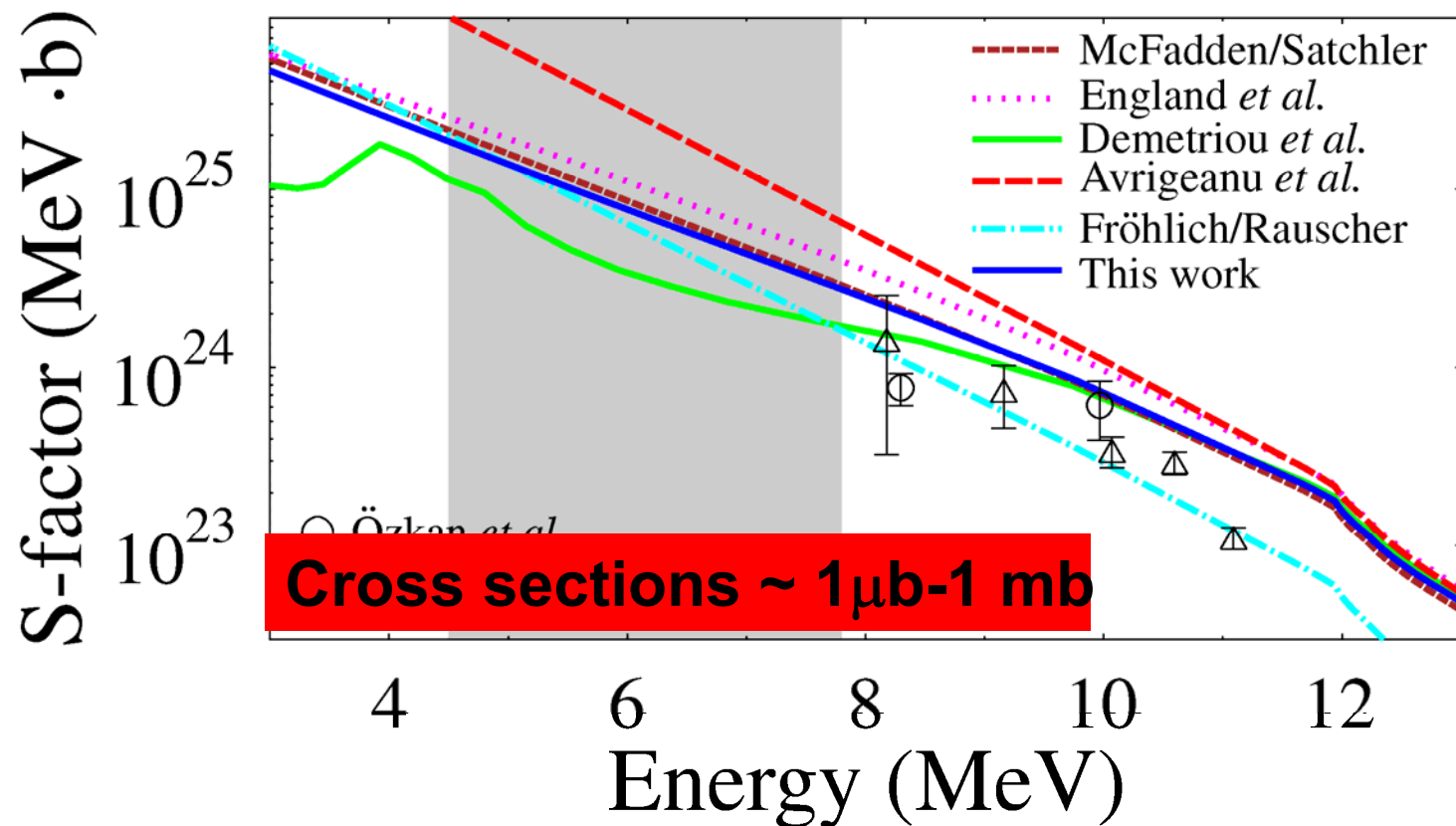
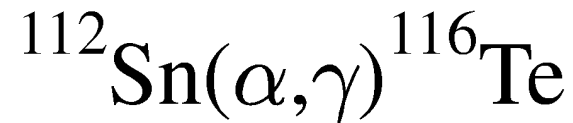
# Proton capture on $^{92}\text{Mo}$ at HORUS



# Predicted $^{92}\text{Mo}(p,\gamma)$ S-Factor



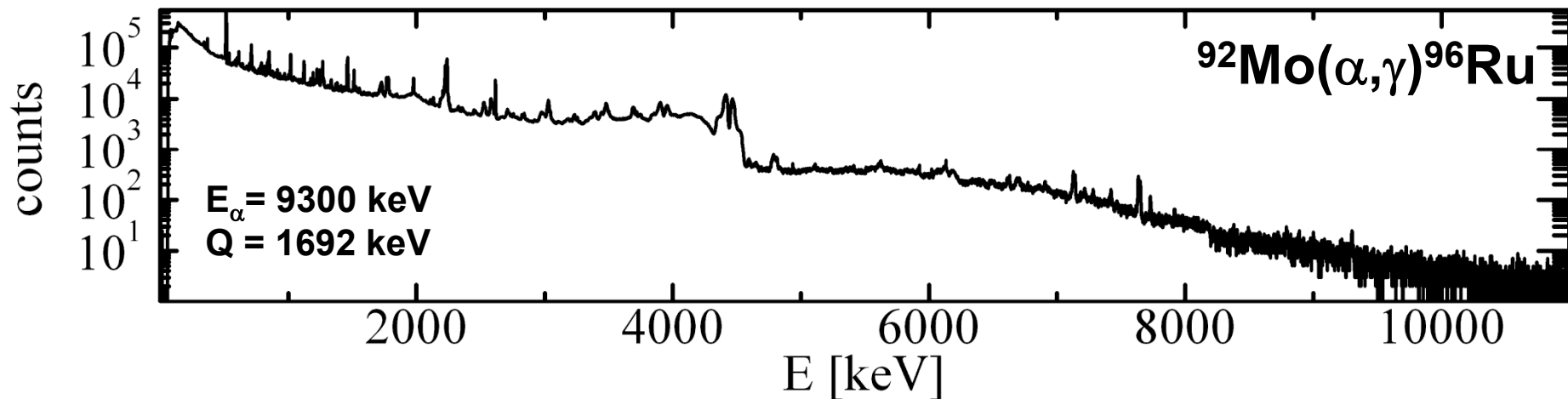
# $\alpha$ - capture cross section



*N. Özkan et al., PRC 75 (2007) 025801*

Overview: *P. Demetriou, C. Grama, S. Goriely, NPA 707 (2002) 253*

# In-beam $\alpha$ – capture experiment at HORUS



## Many background reactions:

### Target chamber

- $^{27}\text{Al}(\alpha,n)^{30}\text{P}$
- $^{27}\text{Al}(\alpha,p)^{30}\text{Si}$
- $^{27}\text{Al}(\alpha,\alpha')^{27}\text{Al}$

### Target contaminants

- $^{13}\text{C}(\alpha,n)^{16}\text{O}$
- $^{17}\text{O}(\alpha,n)^{20}\text{Ne}$
- $^{18}\text{O}(\alpha,n)^{21}\text{Ne}$

### Competing reactions

- $^{92}\text{Mo}(\alpha,p)^{95}\text{Tc}$

### Decay of radioactive reaction products



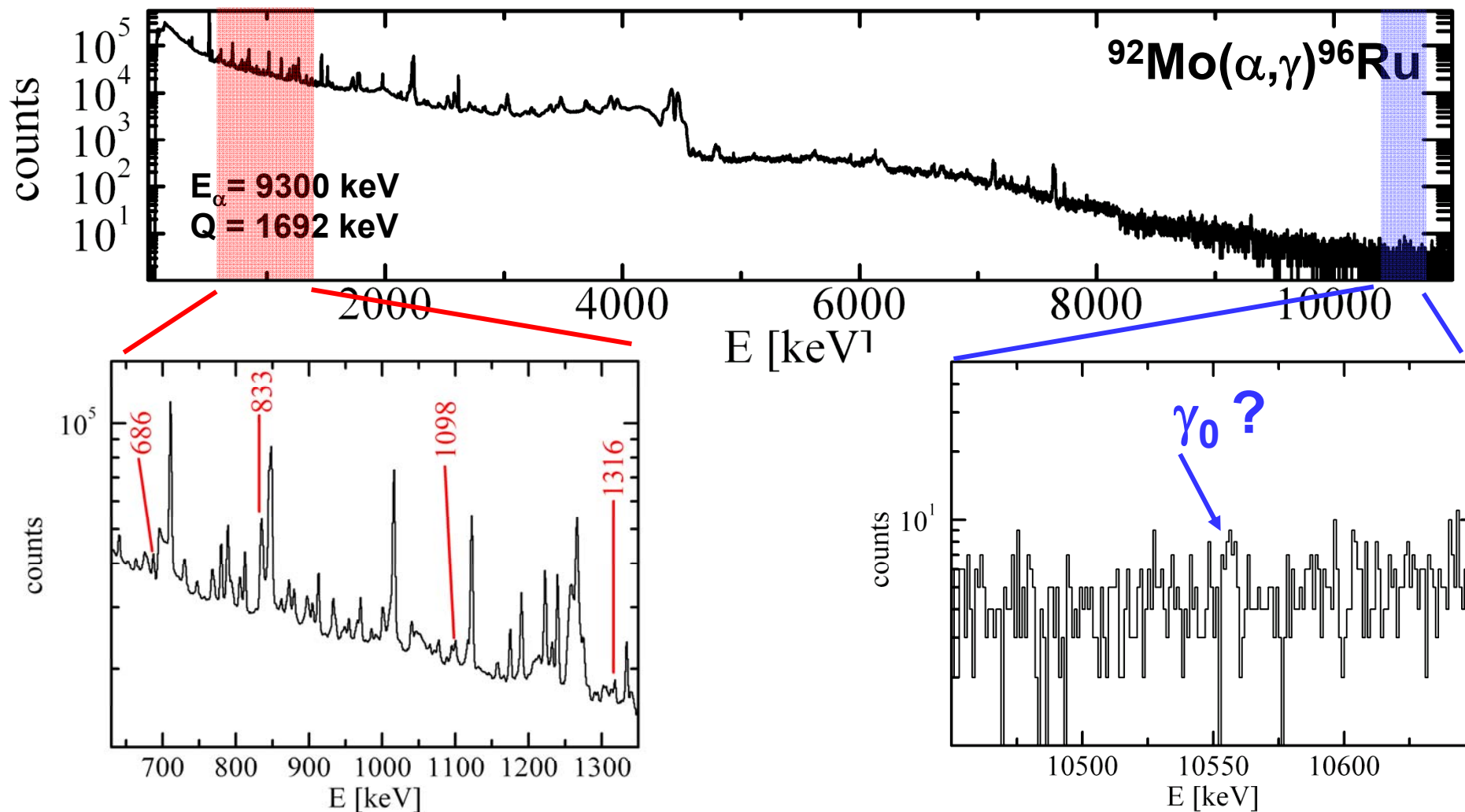
Tantalum coating  
of chamber



High vacuum or  
„Cooling Finger“



# In-beam $\alpha$ – capture experiment at HORUS



**Clear identification of strong low-energetic transitions in  $^{96}\text{Ru}$**

**First indication of  $\gamma_0$**

**PRELIMINARY (experiment in February 2009)**

# $\alpha$ - capture cross sections

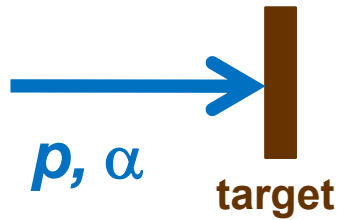
Due to Coulomb barrier ( $\alpha, \gamma$ ) experiments become more and more difficult for heavier nuclei.

Tb 145 30 s $\beta^+$ $\gamma$ 258; 988; 537... m; g	Tb 146 23 s $\beta^+$ $\gamma$ 1579; 1079; 1417...	Tb 147 1.83 m $\beta^+$ $\gamma$ 1398; 1798...	Tb 148 2.2 m $\beta^+$ 2.2... $\gamma$ 784; 632; 882...	Tb 149 4.2 m $\beta^+$ $\alpha$ 3.99 $\gamma$ 796; 165...	Tb 150 3.67 h $\beta^+$ $\gamma$ 638; 650; 827; 438... $\epsilon$ ; $\beta^+$ 3.1; 3.7... $\alpha$ 3.49 $\gamma$ 638; 496...
Gd 144 4.5 m $\beta^+$ 3.3... $\gamma$ 333; 2433; 630; 347...	Gd 145 85 s $\beta^+$ $\gamma$ 721...; 387; 330	Gd 146 48.3 d $\beta^+$ $\gamma$ 155; 116; 115...	Gd 147 38.1 h $\epsilon$ ; $\beta^+$ $\gamma$ 229; 396; 929...	Gd 148 74.6 a $\alpha$ 3.016 $\gamma$ 400	Gd 149 9.28 d $\epsilon$ ; $\alpha$ 3.016 $\gamma$ 150; 299; 347...
Eu 143 2.6 m $\beta^+$ 4.1... $\gamma$ 1107; 1537; 1913; 108; 1805...; g	Eu 144 10.2 s $\beta^+$ 5.2... $\gamma$ 1660; 818...	Eu 145 5.93 d $\epsilon$ $\beta^+$ 1.7... $\gamma$ 894; 1659; 654...	Eu 146 4.51 d $\epsilon$ $\beta^+$ 1.5... $\gamma$ 74...; 633; 63...	Eu 147 24.6 d $\epsilon$ ; $\beta^+$ $\alpha$ 2.91 $\gamma$ 197; 121; 678...	Eu 148 55.6 d $\epsilon$ ; $\beta^+$ $\alpha$ 2.63 $\gamma$ 550; 630; 611...
Sm 142 72.4 m $\epsilon$ $\beta^+$ 1.0 $\gamma$ (679)	Sm 143 65 s $\beta^+$ $\gamma$ (669...)	Sm 144 3.07 $\sigma$ 1.6	Sm 145 340 d $\epsilon$ ; $\gamma$ 61; (492...) $e^-$ $\sigma$ 280	Sm 146 $1.03 \cdot 10^8$ a $\alpha$ 2.455	Sm 147 14.99 $1.06 \cdot 10^{11}$ a $\alpha$ 2.235; $\sigma$ 56 $\sigma_n, \alpha$ 0.0006
Pm 141 20.9 m $\beta^+$ 2.7... $\gamma$ 1223; 886; 194; 1346... g	Pm 142 40.5 s $\beta^+$ 3.8... $\gamma$ 1576...	Pm 143 265 d $\epsilon$ no $\beta^+$ $\gamma$ 742	Pm 144 1.0 a $\epsilon$ ; no $\beta^+$ $\gamma$ 618; 697; 477...	Pm 145 17.7 a $\epsilon$ ; $\alpha$ 2.24 $\gamma$ 72; (67) $e^-$	Pm 146 5.53 a $\epsilon$ ; $\beta^-$ 0.8... $\gamma$ 454; 747; 736... $\sigma$ 8400

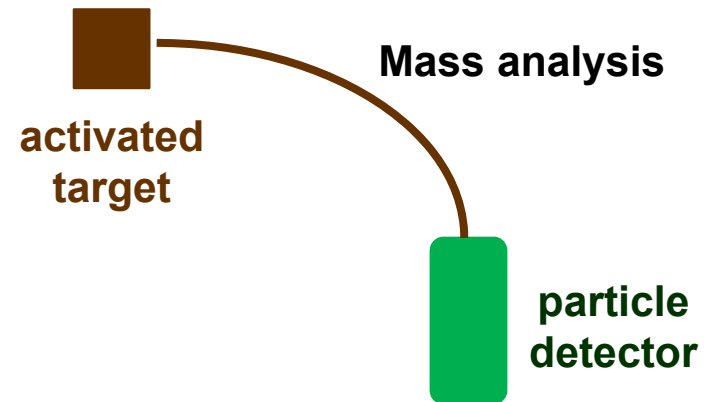
( $\alpha$  spectroscopy: E. Somorjai et al., *Astron. Astrophys.* 333 (1998) 1112)

# Indirect experimental methods III: Mass spectrometry after radiative capture

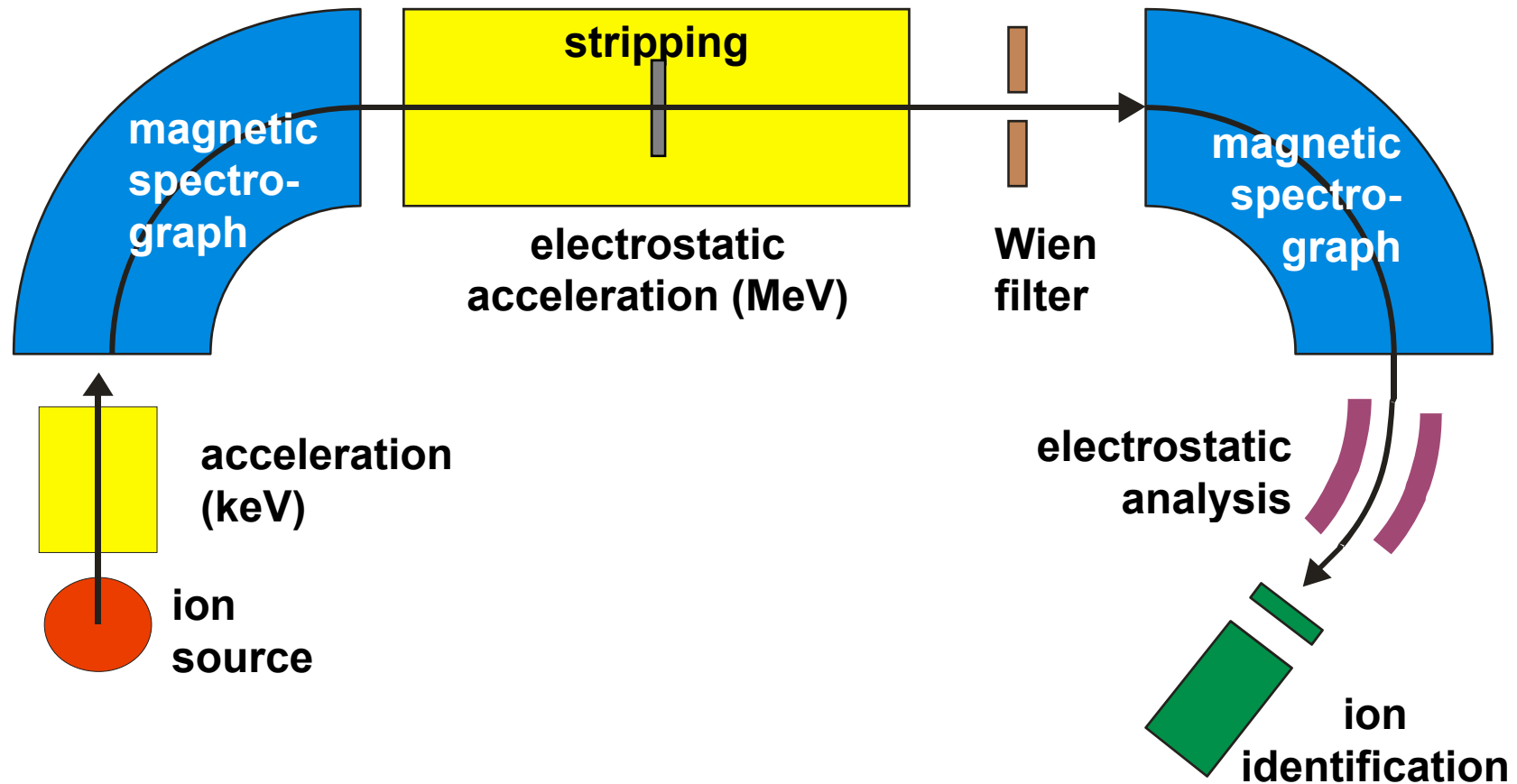
**Step 1**



**Step 2**



# Accelerator Mass Spectrometry (AMS)



**High sensitivity:** isotopic ratios down to  $10^{-15}$ .

**High efficiency:** amounts of  $10^5$  nuclei.

**Separation of isobars possible.**

# Accelerator Mass Spectrometry (AMS)

**Typical application:  
Detection of smallest  
amounts of  $^{14}\text{C}$**

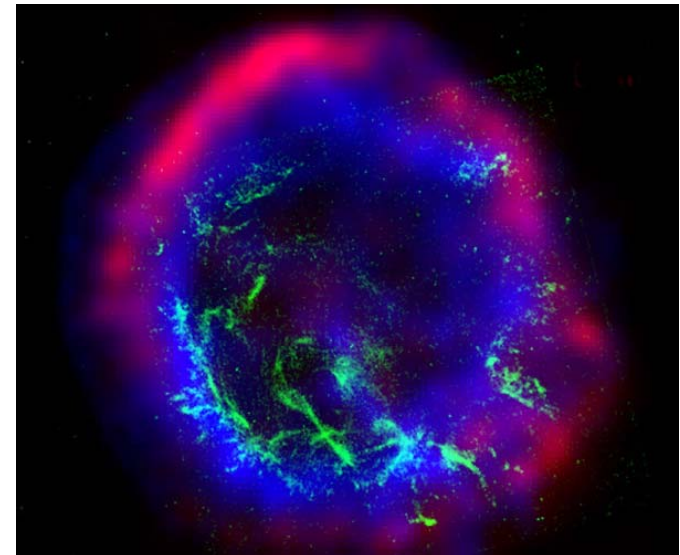


**Further application:  
Detection of other  
cosmogenic nuclides  
(e.g.  $^{10}\text{Be}$ ,  $^{26}\text{Al}$ ,  $^{36}\text{Cl}$ )**



# Accelerator Mass Spectrometry (AMS)

**But as well:  
Detection of other nuclides  
produced in the laboratory  
or by cosmic events**



## **Famous examples:**

*$^{26}\text{Mg}(p,n)^{26}\text{Al}$ : M. Paul et al., PLB 94 (1980) 303*

*$^{40}\text{Ca}(\alpha,\gamma)^{44}\text{Ti}$ : Nassar et al., PRL 96 (2006) 041102*

*$^{60}\text{Fe}$  content in ferromanganese crusts: K. Knie et al., PRL 93 (2004) 17*

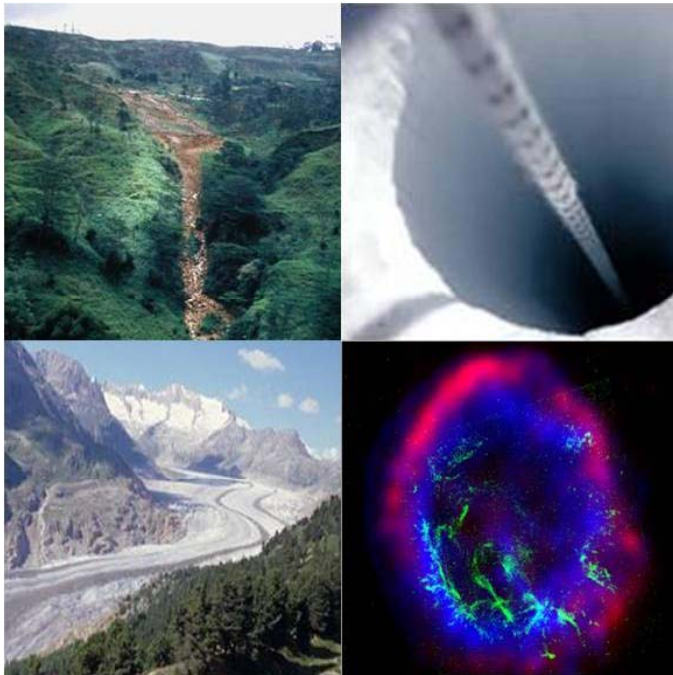
# Accelerator Mass Spectrometry (AMS)

- **Huge demand for AMS in geosciences, archaeology, ...**
- **Existing facilities do not like to test „new“ isotopes**
- **Only a handful „normal“ accelerator laboratories like Notre Dame or MLL Munich are using AMS for non-standard applications**

*(see D. Robertson et al., NIM B 266 (2008) 3481)*

# Accelerator Mass Spectrometry (AMS)

In 2007 the German Funding Agency DFG decided to install a new „high performance“ 6 MV AMS facility in Germany



Application from the University of Cologne:

## Geography

*(U. Radtke)*

## Geology

*(M. Melles, M. Staubwasser)*

## Mineralogy Hannover (now GFZ Potsdam)

*(F. von Blanckenburg)*

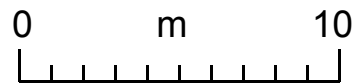
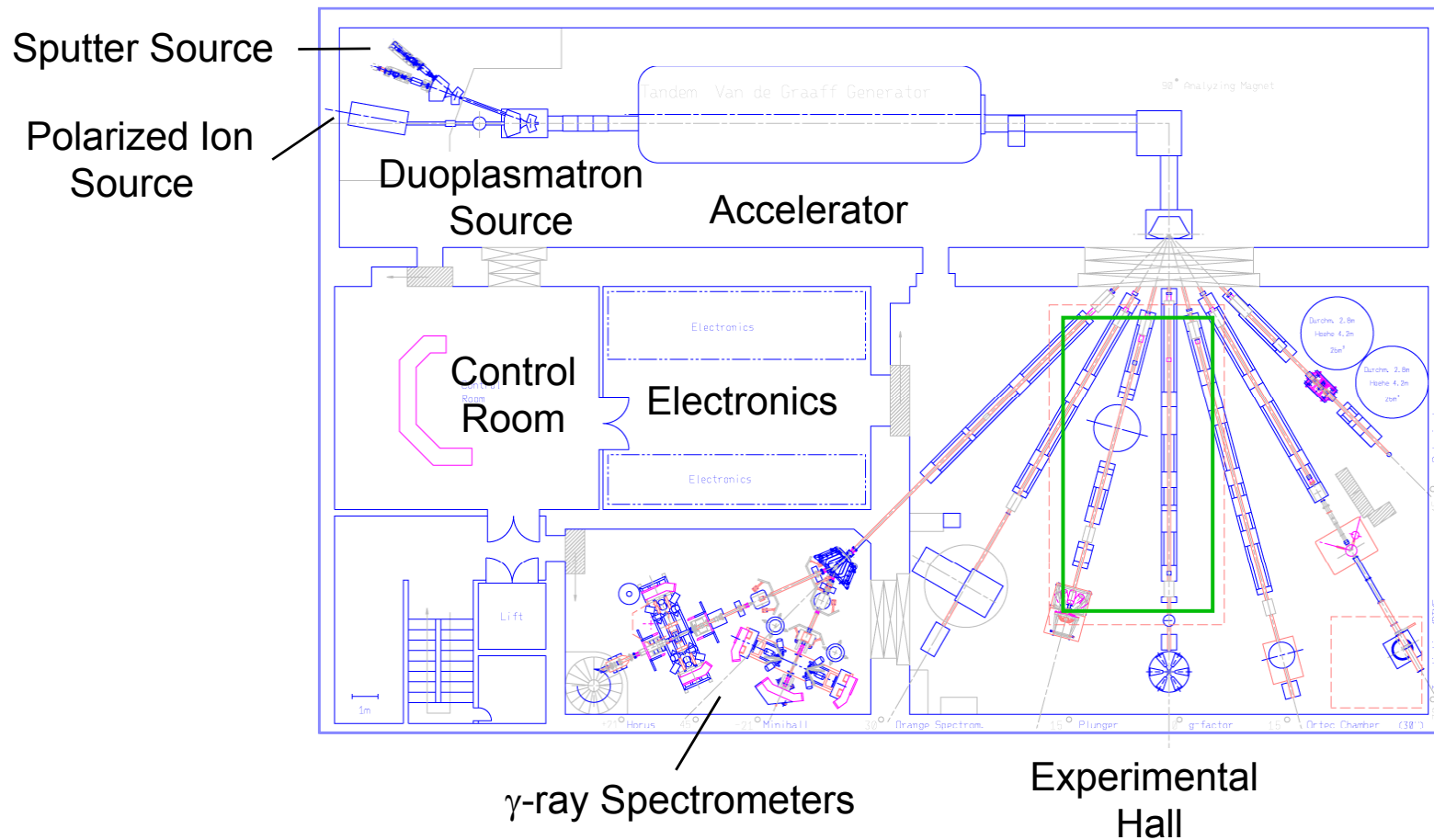
## Nuclear Physics

*(A. Dewald, J. Jolie, A. Zilges)*

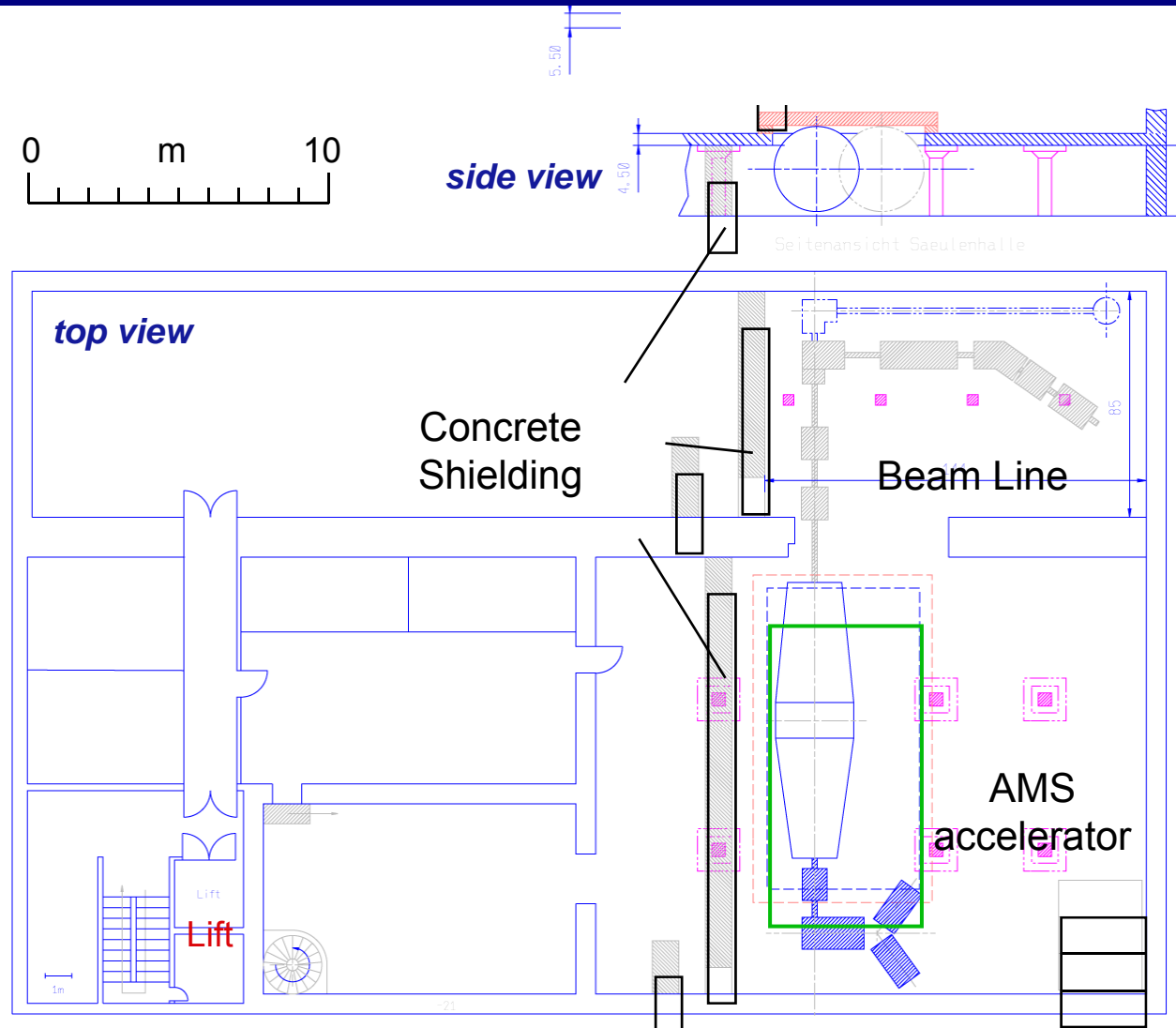
**About 20% of total beamtime reserved  
for developments and nuclear physics !**



# The existing 10 MV Tandem Accelerator (1st basement) Institute for Nuclear Physics, Universität zu Köln



# The new 6 MV Tandem AMS machine (2nd basement)



Funded by **DFG** and



University  
of Cologne

# The new 6 MV Tandetron Accelerator Institute for Nuclear Physics, University of Cologne

Total investment: about 9 M€

Annual running costs: about 0.3 M€

Additional personnel: 2 Professors, 2 Scientists, 3 Technicians, 4 Operators



**AMSCologne**

*A. Dewald, J. Jolie, and A. Zilges, Nuclear Physics News 18 (2008) 26*

Funded by **DFG** and



**University  
of Cologne**

# Facilities at the Universität zu Köln

- The 10 MV Tandem and the 6 MV AMS Cologne run completely independently
- Sample preparation including Nuclear Chemistry on site
- No complicated PAC procedures
- Delivery of AMS tandetron: early 2010
- Commissioning: mid 2010

**COME AND SEE (AND USE) !**



# Experimental techniques to investigate the $p$ process of nucleosynthesis

**M. Büssing, M. Elvers, J. Endres, J. Hasper**

*Institut für Kernphysik, Universität zu Köln*

**L. Kern, S. Müller, D. Savran,**

**A. Sauerwein, V. Simon, K. Sonnabend**

*Institut für Kernphysik, TU Darmstadt*

Supported by **DFG** (ZI 510/5-1) and by BMBF

More information and references: [www.zilges.de](http://www.zilges.de)